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Technical Note

1968-10

Synchronous Satellite
Stationkeeping SimulationM. C. Crocker
E. H. Swenson

14 May 1968

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Lexington, Massachusetts



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MASSACHUSETTS INSTITUTE OF TECHNOLOGY
LINCOLN LABORATORY

SYNCHRONOUS SATELLITE STATIONKEEPING SIMULATION

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Group 63

TECHNICAL NOTE 1968-10

14 MAY 1968

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MASSACHUSETTS

ABSTRACT

A program has been written to simulate the east-west stationkeeping of a synchronous satellite. Different ways of implementing the thrust sequence of rocket motors and solar sails are discussed.

Accepted for the Air Force
Franklin C. Hudson
Chief, Lincoln Laboratory Office

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NOMENCLATURE

a	=	area of the sail
C	=	velocity of light
\vec{F}	=	component of solar force perpendicular to the sail vector
$F_{a\lambda}$	=	average annual force on the satellite by radiation pressure in the longitudinal direction
F_r	=	force on the satellite in the radial direction
F_θ	=	force on the satellite in the colatitude direction
F_λ	=	force on the satellite in the longitudinal direction
$J_{n,m}$	=	coefficients of the tesseral harmonics
$J_{n,o}$	=	coefficients of the zonal harmonics
m	=	mass of the satellite
P_m	=	Legendre polynomials
r	=	distance from the center of the earth to the satellite
R	=	radius of the earth
\vec{R}	=	radius vector
S	=	solar constant = 1.39×10^3 watts/m ²
\vec{S}	=	sail vector
\vec{S}_N	=	normal to the sail
\vec{T}	=	orbital tangent vector
\vec{U}	=	sun vector
V	=	earth's potential
α	=	solar absorptivity
θ	=	colatitude of the satellite
λ	=	longitude of the satellite
λ_{nm}	=	longitude of the n and mth harmonic
μ	=	mass of the earth x gravitational constant
ω	=	earth's rotational rate
I_{sp}	=	specific impulse

SYNCHRONOUS SATELLITE STATIONKEEPING SIMULATION

I. INTRODUCTION

The motion of a satellite with a period of one sidereal day is perturbed by the gravitational attraction of the sun and moon, radiation pressure and the tangential component of the earth's gravitational geopotential. A Synchronous Satellite Simulation program (SSS) was written to study these effects and to determine the feasibility of maintaining a satellite near a specified synchronous position by automatically firing rocket motors and by solar sailing. The different parts of the SSS program are described in the Appendices.

II. UNDERLYING THEORY TO THE COMPUTER SIMULATION

A. Forces Acting On the Satellite

The influence of the gravitational attraction of the sun and moon contribute to a small diurnal oscillation in the longitude of the satellite of the order of 0.5 degrees¹ at the most. A typical daily longitude variation has been calculated by Molitor and Kaplan² and is shown in Fig. 1. However, the sun and moon's gravitational attraction does cause the orbit normal to precess³ so that the inclination changes by about 0.85 degrees per year.

Radiation pressure will change the eccentricity of the orbit a small amount for satellites with ordinary area to mass ratios that are symmetric about a plane that contains the center of mass of the satellite and is perpendicular to the equatorial plane. By making the satellite not symmetric, it is possible to East-West stationkeep a synchronous satellite by radiation pressure forces. Methods for doing this are presented in a paper by Black et al⁴.

B. Orbital Mechanics

The motion of a satellite is determined primarily by the earth's gravitational potential function

$$V = -\frac{\mu m}{r} \left[1 - \sum_{n=2}^{\infty} J_{n,0} \left(\frac{R}{r} \right)^2 P_n^0(\cos \theta) + \sum_{n=2}^{\infty} \sum_{m=1}^{\infty} J_{n,m} \left(\frac{R}{r} \right)^n P_n^m(\cos \theta) \cos(\lambda - \lambda_{n,m}) \right] \quad (1)$$

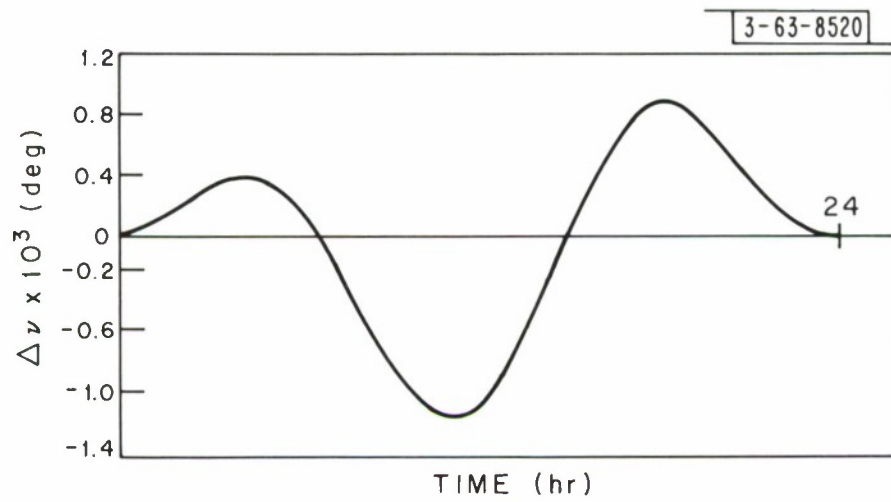


Fig. 1. Typical daily longitude variation of a 24-hour satellite due to solar and lunar perturbations.

For the purposes of studying east-west stationkeeping the effect of the gravitational attraction of the moon, sun and planets was neglected. The SSS program has the option of neglecting terms in the potential function greater than order two or three. The inaccuracies in measuring and producing surfaces with specified reflective properties produce uncertainties in the forces acting on the satellite which mask the above effects.

The motion of a 24-hour satellite is best studied by using an earth fixed system of coordinates rotating at the earth's angular rate ω .^{5,7} The kinetic energy of the satellite expressed in this rotating system is

$$T = \frac{1}{2} m [\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta (\dot{\lambda} + \omega)^2] \quad (2)$$

Let us denote the forces on the satellite produced by rocket thrusters or radiation pressure devices as F_r , F_θ , F_λ . Now by using Lagrange's equations and retaining terms to the second order, the following equations of motion result.

$$\begin{aligned} \ddot{r} - r \dot{\theta}^2 - r \sin^2 \theta (\dot{\lambda} + \omega)^2 &= -\mu/r^2 + 3\mu J_{20} R^2 (3 \cos^2 \theta - 1)/2r^4 \\ &+ 9\mu J_{22} R^2 \sin^2 \theta \cos 2(\lambda - \lambda_{22})/r^4 + F_r/m \end{aligned} \quad (3)$$

$$\begin{aligned} \ddot{\theta} &= \frac{3\mu J_{20} R^2}{2r^5} \sin 2\theta - \frac{3\mu J_{22} R^2}{r^5} \sin 2\theta \cos 2(\lambda - \lambda_{22}) \\ &+ \frac{(\dot{\lambda} + \omega)^2}{2} \sin 2\theta - 2\dot{r}\dot{\theta}/r + F_\theta/mr \end{aligned} \quad (4)$$

$$\begin{aligned} \ddot{\lambda} &= \frac{6\mu J_{22} R^2}{r^5} \sin 2(\lambda - \lambda_{22}) - 2 \cot \theta \dot{\theta} (\dot{\lambda} + \omega) + F_\lambda/mr \sin \theta \\ &- \frac{2\dot{r}(\dot{\lambda} + \omega)}{r} \end{aligned} \quad (5)$$

When the force terms are set to zero F_θ , F_r , and F_λ the program has been found to simulate the motion of Syncom II to a high degree of accuracy as shown in Fig. 2.⁶ The discontinuity in the curves is due to a rocket motor firing.

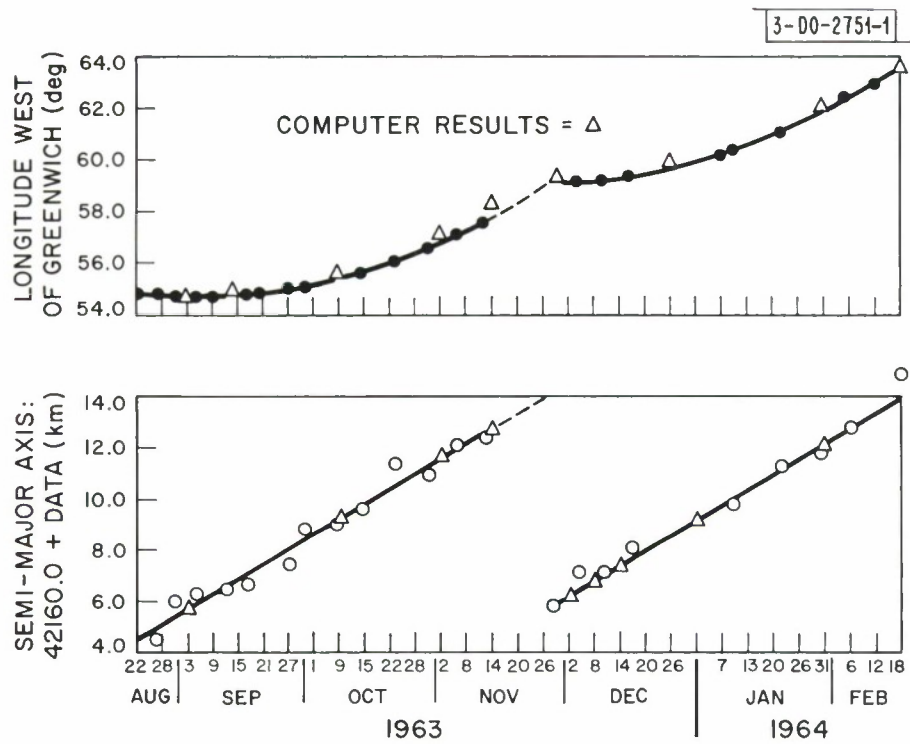


Fig. 2. SYNCOM II drift.

C. Longitudinal Perturbation Force

The longitudinal force per unit mass on a satellite due to the earth's geopotential is given by taking the gradient of the potential

$$f_{\lambda}/m = - \nabla_{\lambda} V/m = \frac{1}{mr \sin \theta} \frac{\partial V}{\partial \lambda} = \frac{6J_{22} R^2 \mu}{r^4} \sin \theta \sin 2(\lambda - \lambda_{22}) \quad (6)$$

For an equatorial orbit $\theta = 90^\circ$ and

$$f_{\lambda}/m = - 5.568 \times 10^{-8} \sin 2(\lambda - \lambda_{22}) \text{ newtons/kilogram} \quad (7)$$

The stable equilibrium points are located at the positions $(\lambda - \lambda_{22}) = 90^\circ$ and 270° .

If a satellite were placed at these points, the satellite would remain there or oscillate about them with a small amplitude. No stationkeeping apparatus would be required if the satellite were placed at these points and the portion of the earth "visible" to the satellite was satisfactory. The stable equilibrium points are located off the western coast of South America and over the Indian Ocean. However, if the satellite is used for communications between the United States and Europe or the United States and the Far East the satellite will require stationkeeping apparatus. The visibilities of satellites placed for communication between the United States and Europe and between the United States and the Far East are shown in Fig. 3.

III. STATIONKEEPING BY SOLAR SAILING

The instantaneous radiation pressure forces, on the sail of a satellite, which has its plane in the north-south direction are given by the following vector equations, assuming there is no angular dependence of the solar absorptivity. These equations hold for any orbit.

$$F_{\lambda} = \frac{Sa}{c} (h(2 - \alpha) (\vec{S}_N \cdot \vec{U})^2 - |\vec{S}_N \cdot \vec{T}| - s\alpha |\vec{S}_N \cdot \vec{U}| (\vec{U} \cdot \vec{S}) \sin [\cos^{-1} (-\vec{S}_N \cdot \vec{T})]) \quad (8)$$

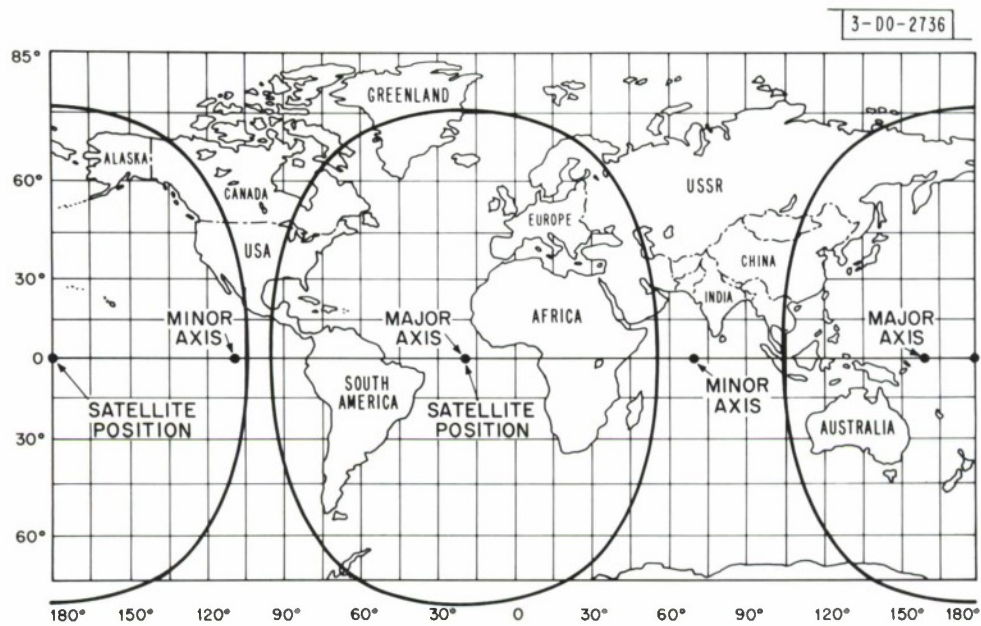


Fig. 3. Visibilities of satellites placed at stable equilibrium points.

$$F_r = \frac{Sa}{c} [g(2 - \alpha) (\vec{S}_N \cdot \vec{U})^2 \sin [\cos^{-1}(-\vec{S}_N \cdot \vec{T})] - f\alpha |\vec{S}_N - \vec{U}| |\vec{S}_N - \vec{T}| (\vec{U} - \vec{S})] \quad (9)$$

$$F_\theta = -\frac{Sa}{c} \alpha \cdot |\vec{S}_N \cdot \vec{U}| U_z \quad (10)$$

where h, s, g, f are +1 with the following exceptions:

$$\begin{aligned} h &= -1 & \text{when } (\vec{F} \cdot \vec{T}) &\leq 0 \\ s &= -1 & \text{when } (\vec{S} \cdot \vec{T}) &\leq 0 \\ g &= -1 & \text{when } (\vec{F} \cdot \vec{R}) &\leq 0 \\ f &= -1 & \text{when } (\vec{S} \cdot \vec{R}) &\leq 0 \end{aligned}$$

A. Sail Type I

The simplest type of sail for a satellite that has one side oriented to the earth and one side oriented toward the north direction, is a sail whose plane is coincident with the radius to the satellite from the center of the earth and the north-south direction. (See Fig. 4) This sail has a low solar absorptivity on one side and a high solar absorptivity on the other. The thermal emissivity of both sides is approximately the same. Under these conditions the force resultant from thermal radiation from both sides of the sail cancels for a thin sail. The net force along the orbit is then due only to solar radiation pressure. The average annual force in the longitudinal direction is given by

$$F_{a\lambda} = \frac{\alpha_2 - \alpha_1}{4} \frac{Sa}{c} < \cos i > \quad (11)$$

where $< \cos i >$ is average annual effect of the sun's movement above and below the orbital plane of the satellite. If one assumes that the earth's orbit about the sun is circular and an equatorial orbit for the satellite $< \cos i > \approx 0.96$ and α_2 is the solar absorptivity of the sail facing in the direction of orbital

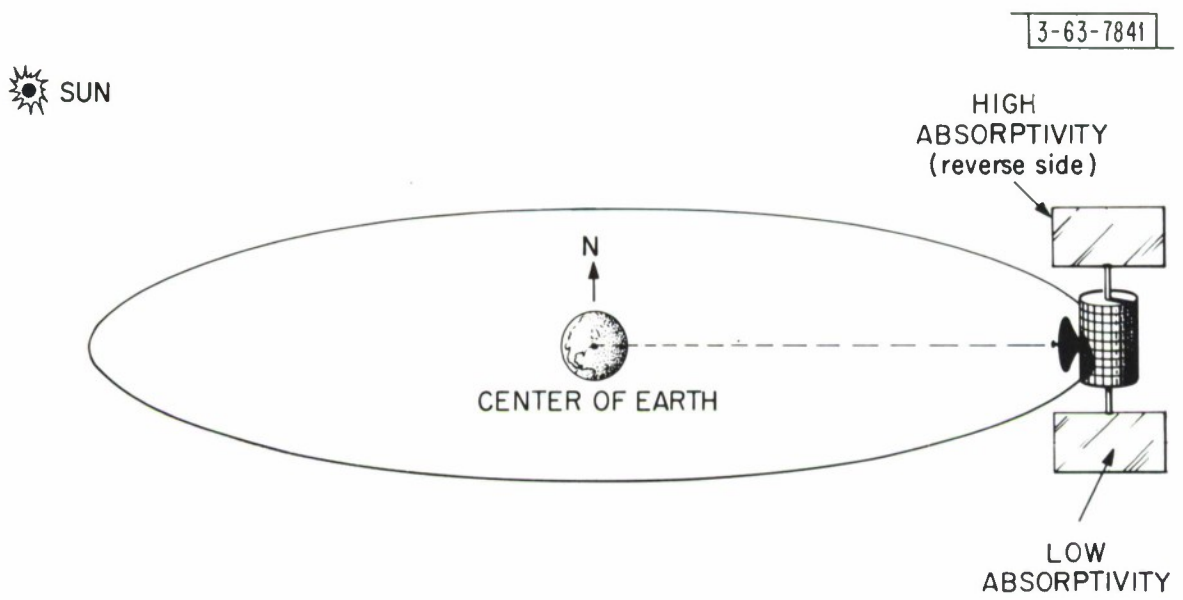


Fig. 4. Solar sail Type I.

motion and α_1 is solar absorptivity of the opposite face. This type of sail does not lend itself to being adapted to the stationkeeping methods to be discussed in Section IV.

B. Sail Type II

The second type of sail is similar to the first except both sides of the sail have the same highly reflective surface and can be rotated along the earth satellite line so that the plane of the sail goes from the north-south to east-west direction. The sail is left in the north-south position for half the orbit starting (or ending) with the sun-earth line and the east-west position for the other half of the orbit. (See Fig. 5) The average annual force in the longitude direction is

$$F_{a\lambda} = \left[\pm \left(\frac{2 - \alpha}{4} \frac{Sa}{c} \right) \langle \cos i \rangle - \alpha \frac{Sa}{2\pi c} \langle \sin 2i \rangle \right] \quad (12)$$

The positive sign is used when the north-south position is taken for half the orbit starting with the earth sun line and the negative sign is used when the north-south position ends with the earth sun line. This type of sail is more complicated than the first type because it requires a rotating mechanism and sensor logic to determine the location of the earth sun line. It does require less sail area than the first type. It is also adaptable to the previously mentioned methods of thrusting by rotating the sail into or out of position with respect to the earth-sun line.

C. Sail Type III

The third type of sail has both sides highly reflective. It rotates about the north-south direction continuously with a period of 12 hours. (See Fig. 6) The average annual force in longitudinal direction is

$$F_{a\lambda} = \frac{Sa}{\pi c} \frac{2}{3} + 2(1 - \alpha) \sin 2\varphi \langle \cos i \rangle \quad (13)$$

where φ is the angle between the plane of the sail and the projection of the earth sun line on the equatorial plane when the satellite is crossing the projection of the earth sun line on the equatorial plane. This type of sail requires the least



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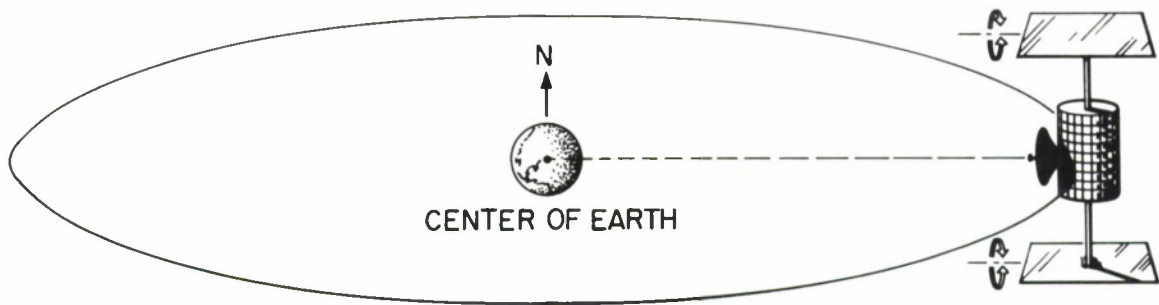


Fig. 5. Solar sail Type II.



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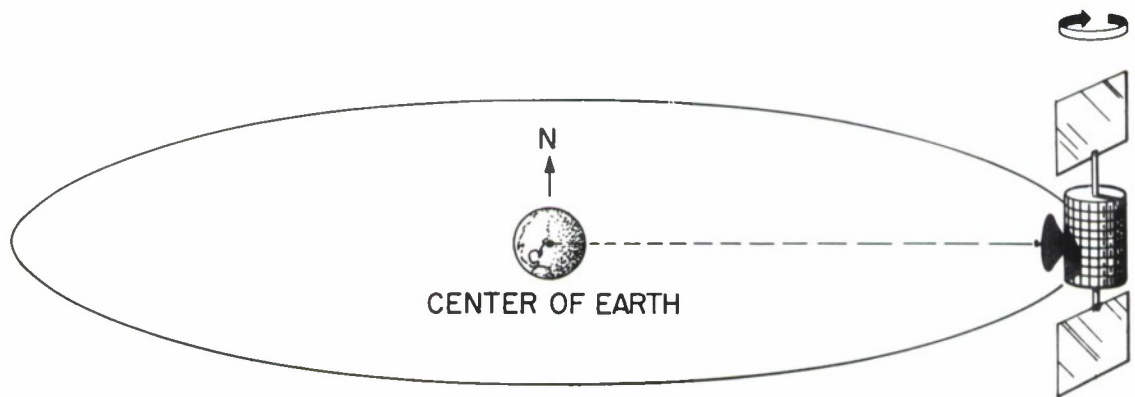


Fig. 6. Solar sail Type III.

amount of sail area for stationkeeping. If one sets $\varphi = 45^\circ$ and assumes a solar absorptivity of 0.9 for one side of the Type I sail and 0.05 for the other types of sails, the relative area needed to produce the same average annual force in the longitudinal direction for Type I, Type II and Type III sails is in the ratio of 3.84 : 1.70 : 1 respectively. A typical satellite with a mass of 100 Kg would require a Type III sail with an area of 1.47 square meters to stationkeep at the position of highest longitudinal force due to the earth's geopotential. By varying the value of φ it is possible to vary the longitudinal force along or against the orbital motion. Therefore, this type of sail is adaptable to any method of automatic stationkeeping.

D. Results of Computer Simulations

The computer simulation of stationkeeping a synchronous satellite by solar sailing revealed little difference between the three types of sail considered for stationkeeping. A typical stationkeeping simulation is shown in Fig. 7. In this simulation sail Type III was used with $\varphi = -45^\circ$ and $\alpha = .1$. The area of the sail was 1.35 square meters for a 100 Kg satellite. The satellite was stationkept near a position at which a maximum longitudinal force occurs. (296° E longitude)

IV. AUTOMATIC STATIONKEEPING

A. Stationkeeping by Constant Force

Different methods of applying a tangential force to counteract the longitudinal perturbational force were simulated by the program. The simplest of these is to fire a rocket motor with the same value of longitudinal force that is acting upon the satellite. However, in practice the force cannot be exactly matched for each longitudinal position of the satellite. The method is limited if there is no control over the force level, once the satellite is placed in a longitudinal position within the longitude regions $45^\circ < \lambda - \lambda_{22} < 135^\circ$ and $225^\circ < \lambda - \lambda_{22} < 315^\circ$. If the force level is not exactly matched for a particular longitudinal position the satellite will undergo an oscillation about the longitude for which it is exactly matched. A simulation of an unmatched force level was done by the SSS program and is shown in Fig. 8. The satellite was initially at 291° longitude and the force level was equal to the longitudinal force at 283.5° longitude.

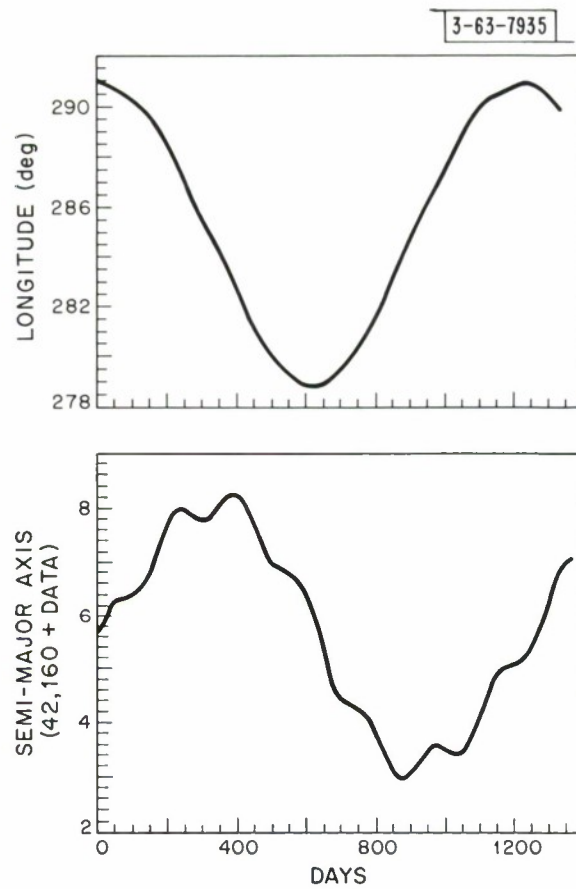


Fig. 7. Variations in semi-major axis and longitude versus time.

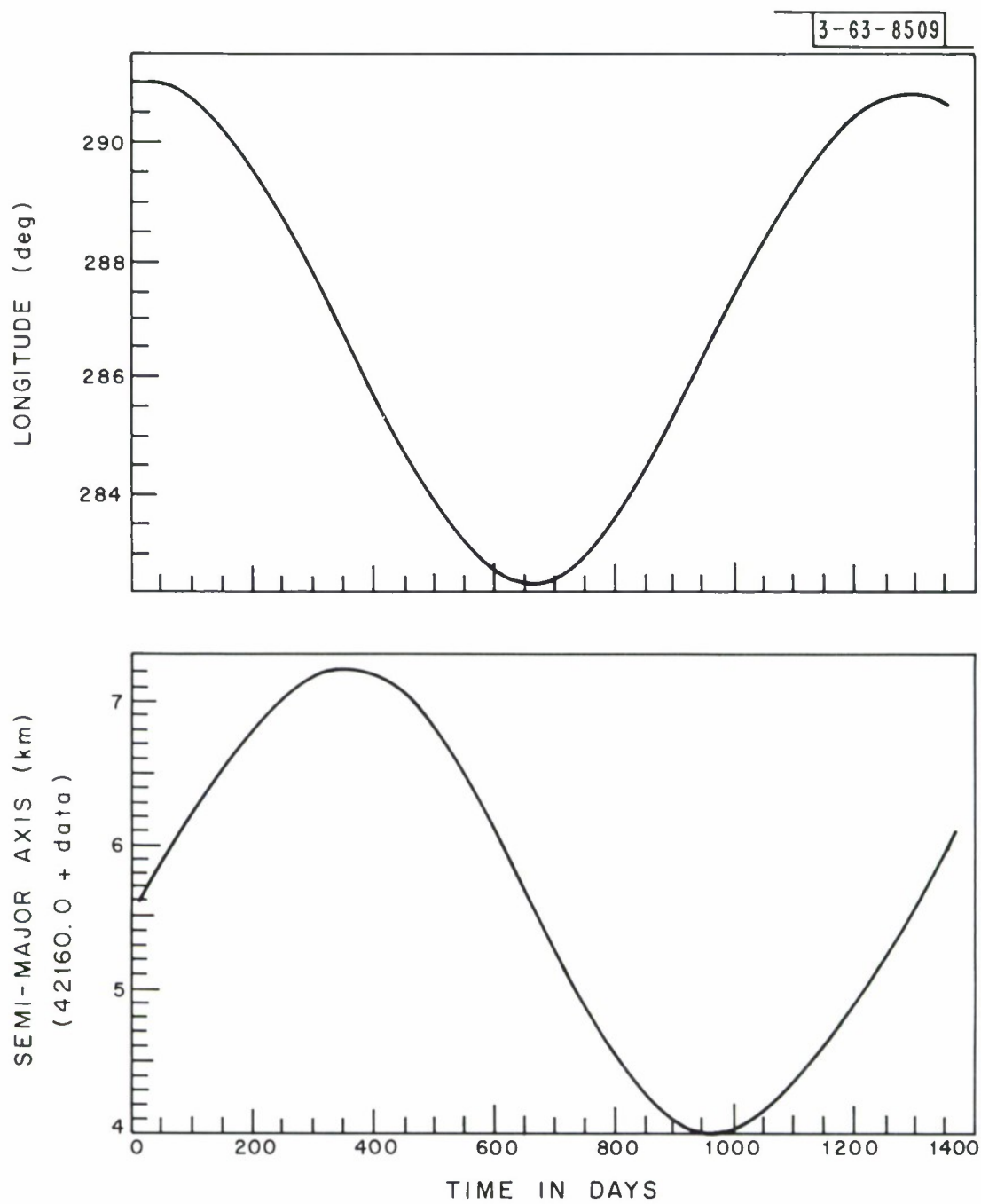


Fig. 8. Simulation of an unmatched force level.

B. Automatic Stationkeeping - Unidirectional Firing

Automatic stationkeeping was proposed by W. E. Morrow⁸ in a memo dated 1 July 1965. A block diagram of the logic required by the satellite for internally computing its position and rocket firing times is shown in Fig. 9. This description does not describe the LES-6 stationkeeping logic.

It is the purpose of the coincidence logic to provide a pulse derived from fan beam sensors at a predetermined satellite position with respect to the earth and sun. The satellite may be assumed spinning so that the output from the sensors are in the form of pulses or for non-spinning satellite the output from the sensors may be chopped to produce the pulses. This coincidence pulse permits comparison between a running clock and a fixed clock in the decision logic. Almost any position in the orbit can be used for establishing the satellite's position where the optical sensors can view an illuminated portion of the earth. For definiteness let us pick the 6 a.m. position shown in Fig. 10.

The decision logic determines whether the satellite is early or late in arriving at the "6 a.m. position" each day. The device compares the coincidence signal with the output signal from the clock logic.

If the signal from the coincidence logic occurs in position A of Fig. 11, the satellite is early for the three days shown and no signal is sent to the thrust logic. If a signal from the coincidence logic falls in position B then a signal is sent to the thrust logic.

The thrust logic's purpose is to fire the rocket motor in the direction of the orbit tangent for a specified length of time near the 6 a.m. position. This is repeated daily until the satellite is "on station". The thrusting is then stopped until another signal is received.

The clock must have an accuracy of $5 \text{ in } 10^{10} / \text{day}$ for 5 years and put out a signal (alarm) at 10 hr 12 min GMT each day. This signal is sent to the decision logic for further processing.

The purpose of the clock logic is to apply a daily correction to the clock to allow for the fact that the sun is ahead of or behind the clock according to

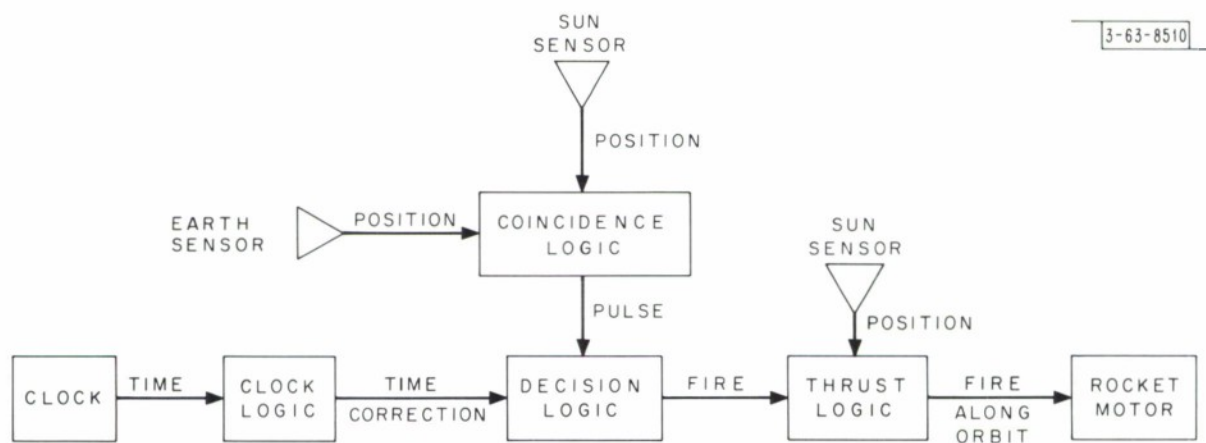


Fig. 9. Satellite logic for computing position and rocket firing times.

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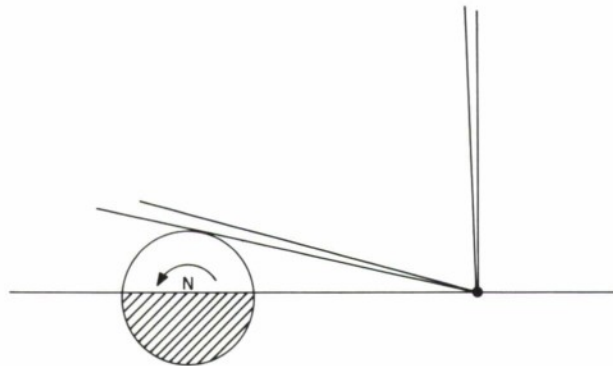


Fig. 10. 6 a.m. position of satellite in orbit.

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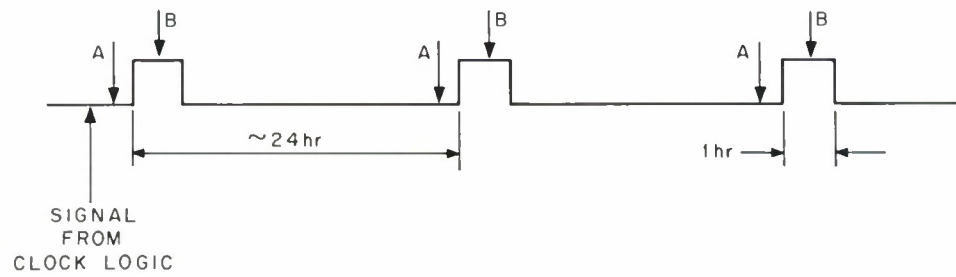


Fig. 11. Signal from clock logic.

the time of year. The correction that must be added to the time is shown in Fig. 12. The curve shown in Fig. 12 has to be approximated by a function that is easy to program into the satellite logic. The errors associated with the degree of exactness to which the clock logic will correct for the difference between apparent solar and solar mean time were accounted for from tables furnished by A. Braga-Illa for LES-6.

As an example of how the correction is applied, let us consider the correction at 15 January 1967. To do an exact calculation of the time to be added we would have to go through an iterative process. However, for our purposes the correction may be taken as the correction that occurs on January 15.0 which is 9.0 min. Therefore, the clock logic is to put out a signal to the decision logic which starts at 10 hr 21 min GMT and lasts for one hour. When the sun is ahead of the clock the correction has to be subtracted from 24 hours and the result added to the time of the previous days output.

The method is a little cumbersome in logic design. A better way to do it would be to shift 0 time base line so that all corrections become additions.

It is possible to do without the clock logic by using an optical sensor that senses the declination of the sun and relative position of the earth satellite line at the same time to correct for the actual longitudinal motion of the sun relative to the "mean" sun in the celestial sphere. The sensor can be visualized as having a cluster of pencil beams arranged in the familiar analemma found on some globes.⁹ A sensor of this type is being constructed by C. Burrowes for use in LES-6.

C. Effects of Sensor and Clock Errors

There are errors associated with the measurement of longitude due to the beamwidth of the sensors and the degree of exactness to which the clock logic will correct for the difference between apparent solar and solar mean time. There are also errors induced because the LES-6 satellite spin axis is not aligned perfectly in the north-south direction. The source of error has been investigated by B. J. Moriarty.¹⁰ The shape of the probability density function of the sum of the errors is not known. Therefore, a uniform probability

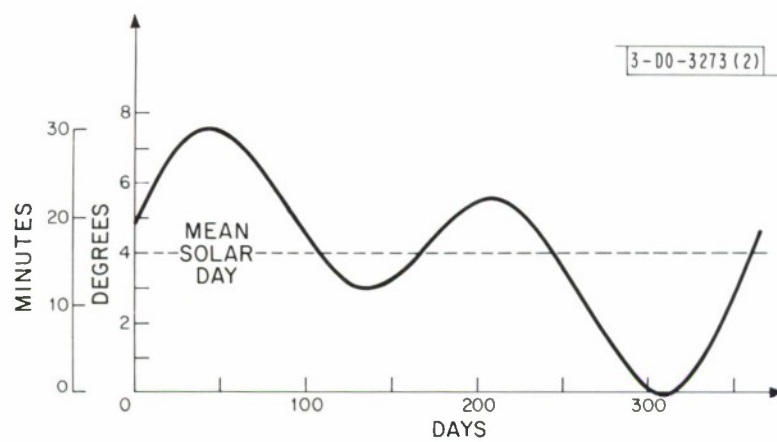


Fig. 12. Ephemeris correction vs. time of the year.

function with a width of one degree of longitude was assumed for use in the SSS program. The errors associated with the analemma sensor for LES-6 were separately accounted for from tables furnished by C. Burrowes.

The influence that the errors have on the motion of the satellite is shown in Figs. 13 and 14. Fig. 13 is the result of an SSS computer simulation of an automatic stationkeeping system using unidirectional firing with no errors present in determining longitude. Fig. 14 is the result of a simulation under the same conditions as the simulation of Fig. 13 except the clock logic errors and sensor errors are introduced. The errors do not significantly affect the range of longitudes over which the satellite may be stationkept. The errors in longitude sensing were less than 0.7 degrees due to clock errors and less than 0.5 degrees due to random sensor errors.

D. Automatic Stationkeeping Unidirectional Firing With Force Cut-Down

The best method of automatic stationkeeping so far proposed is that proposed by Roger Brockett.¹¹ This method uses the least amount of fuel and stationkeeps within the smallest range of longitude. As an example of how this method operates, let us assume the satellite is initially located at a longitude such that the earth's gravitational component along the orbit will cause the longitude of the satellite to decrease in time. The longitude of the satellite is allowed to decrease until it goes past a selected longitude, λ_0 . At this point the rocket motors are fired for a specified length of time such that the impulse given per day is approximately five times the "impulse" given to the satellite by the earth's geopotential. This is the same as the method previously described. However, the longitude is sampled once per day and its value is stored in a memory bank in the satellite. If the longitude measured on a given day is greater than that measured the day before, the rocket impulse given per day is reduced until the satellite again passes the selected longitude, λ_0 , at which time the thrusting is stopped. The amount that the impulse may be reduced is a variable that may be put in the input to the SSS program. A typical example of this stationkeeping method is shown in Fig. 15.

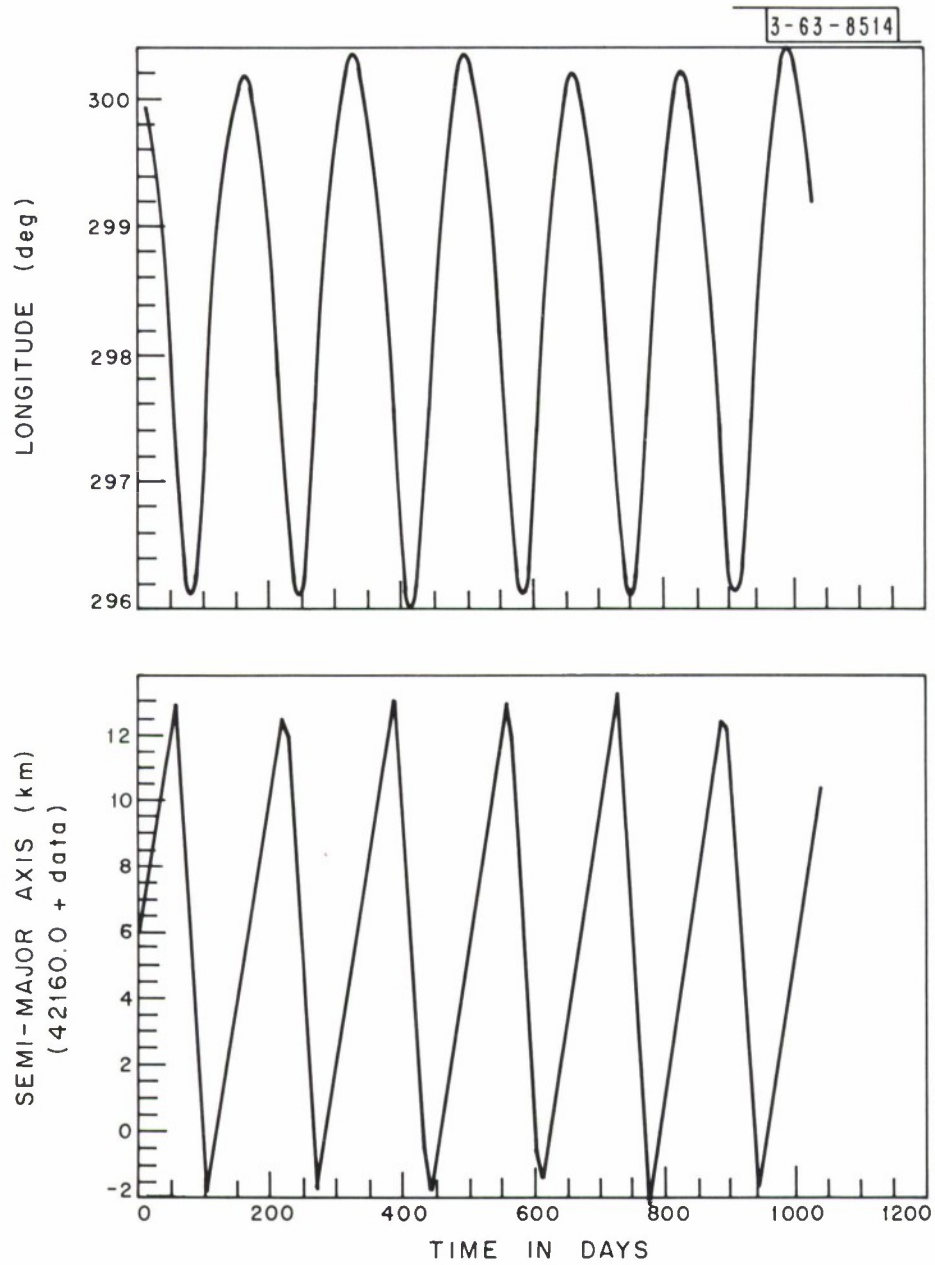


Fig. 13. Results of automatic stationkeeping with unidirectional firing with no errors in determining longitude.

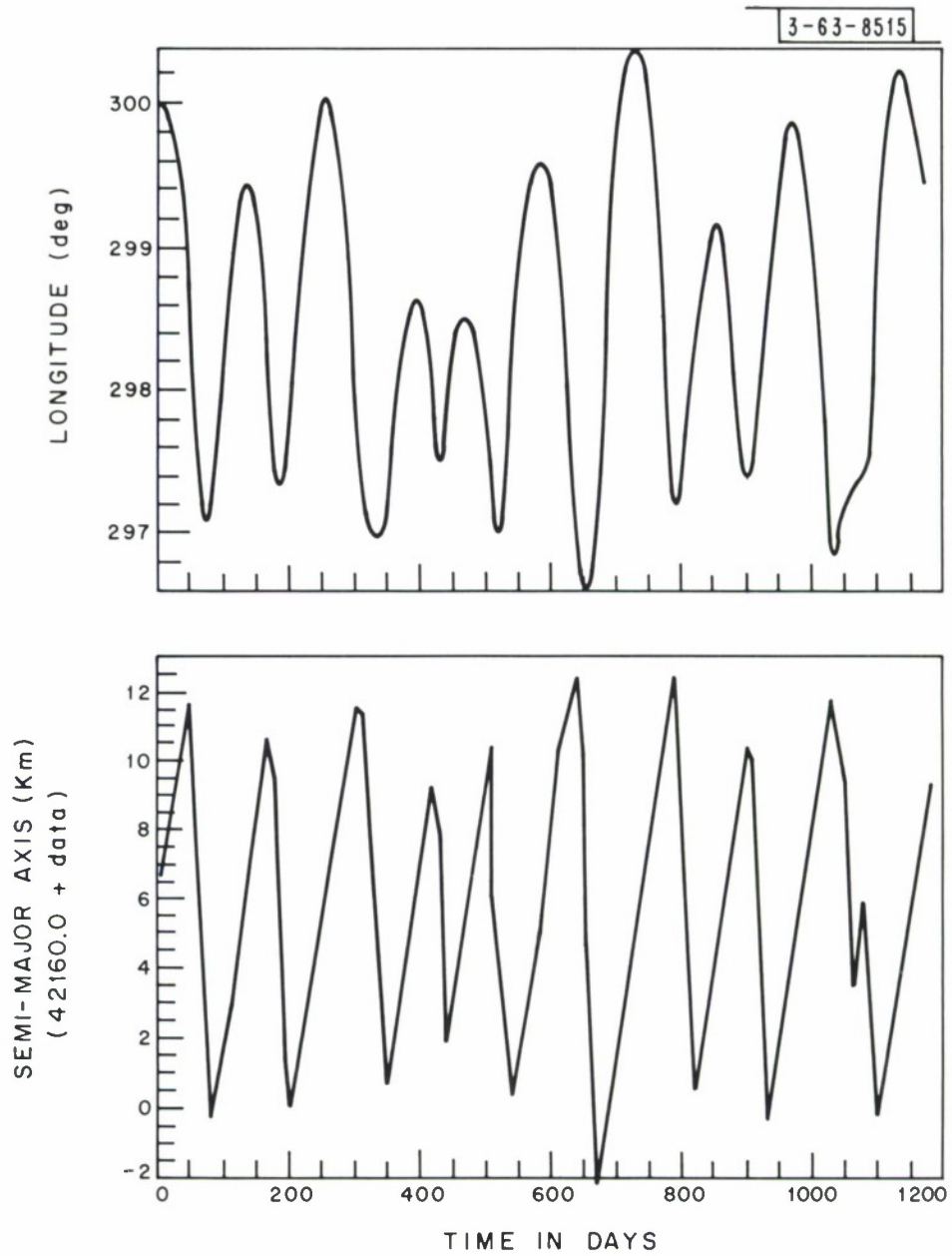


Fig. 14. Results of automatic stationkeeping with unidirectional firing with clock logic error and sensor errors introduced.

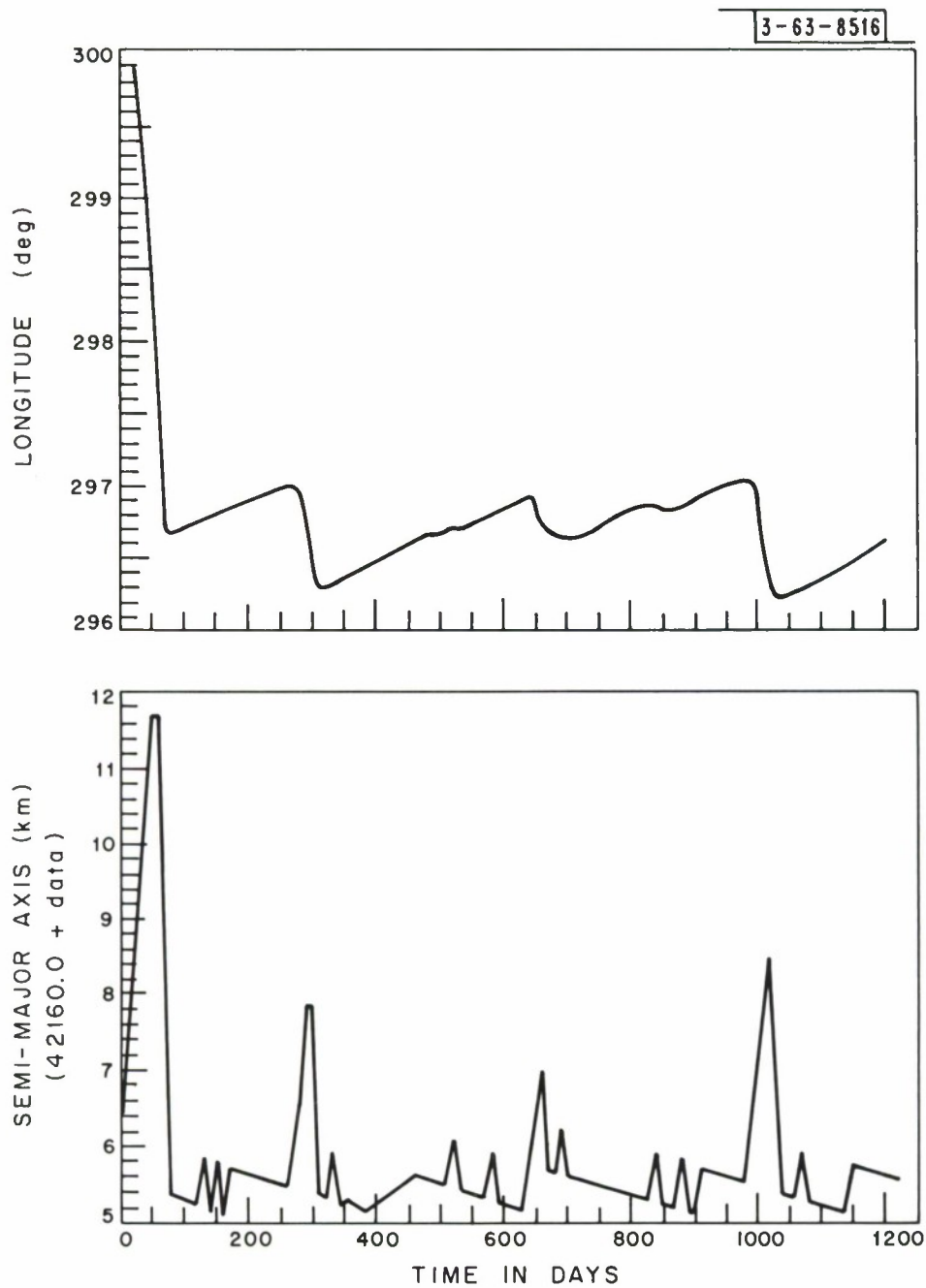


Fig. 15. Results of automatic stationkeeping with unidirectional firing with force cut-down.

E. Automatic Stationkeeping Bidirectional Thrusting

It is possible to stationkeep with a system which thrusts in one direction when the satellite exceeds a given longitude and in the other direction when the satellite is less than another given longitude. In other words, whenever the satellite has drifted out of a given band of longitude, thrusting is done to return the satellite to the band. This method of stationkeeping is not recommended for use with thrust systems that consume fuel at a low specific impulse because a lot more fuel is needed than with the methods previously mentioned. However, it is adaptable to solar sailing techniques where no fuel is used. A simulation of the bidirectional technique is shown in Fig. 16. This simulation was run under the same conditions as the unidirectional firing simulation of Fig. 13. The fuel consumption for stationkeeping a 132 Kg satellite for 1000 days was 3.85 Kg for a rocket system which has an $I_{sp} = 70$ for bidirectional thrusting. Under the same conditions only 0.86 Kg are needed for unidirectional thrusting.

Another method of bidirectional thrusting has been proposed by A. Bragaglia. The logic of this method is explained in Appendix E describing subroutine Newfor of the SSS program. It is well suited for solar sailing techniques. A simulation of this method is shown in Fig. 17.

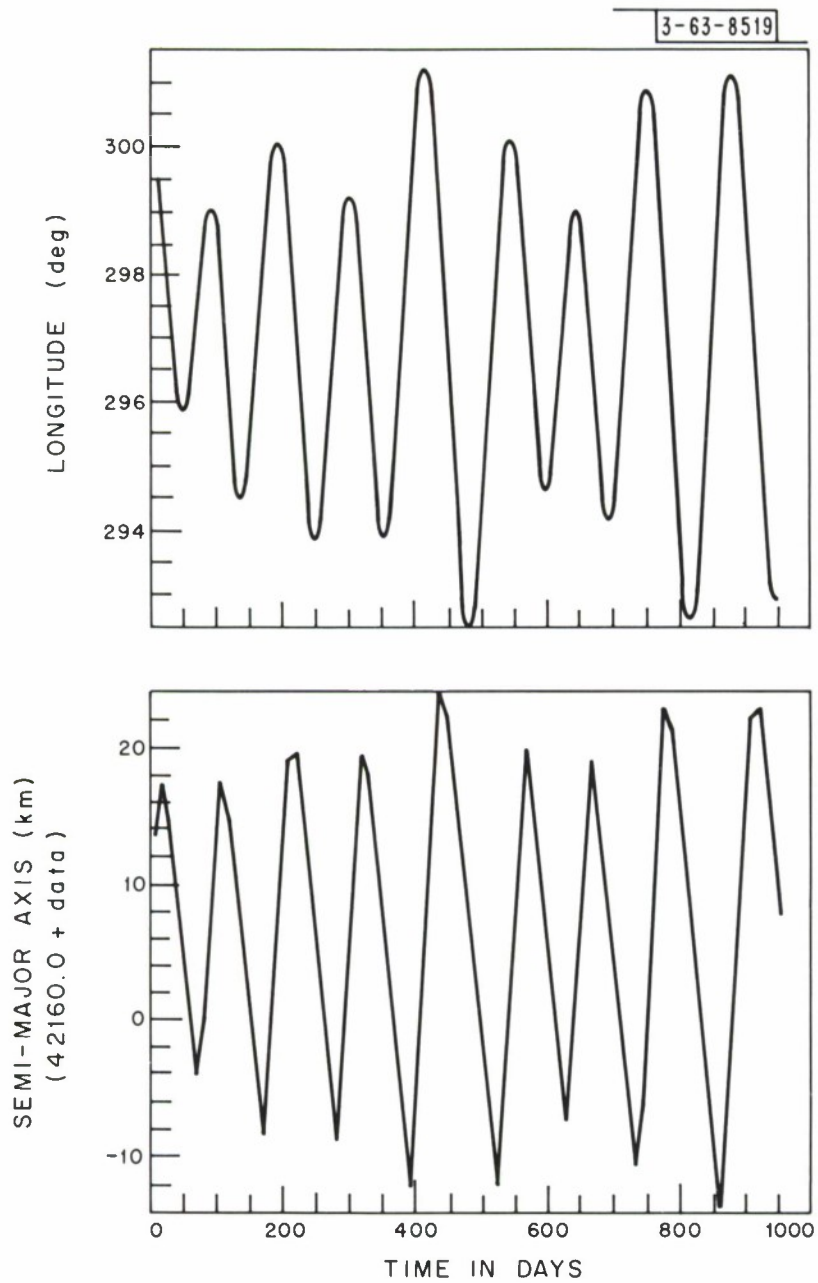


Fig. 16. Results of automatic stationkeeping using bidirectional firing.

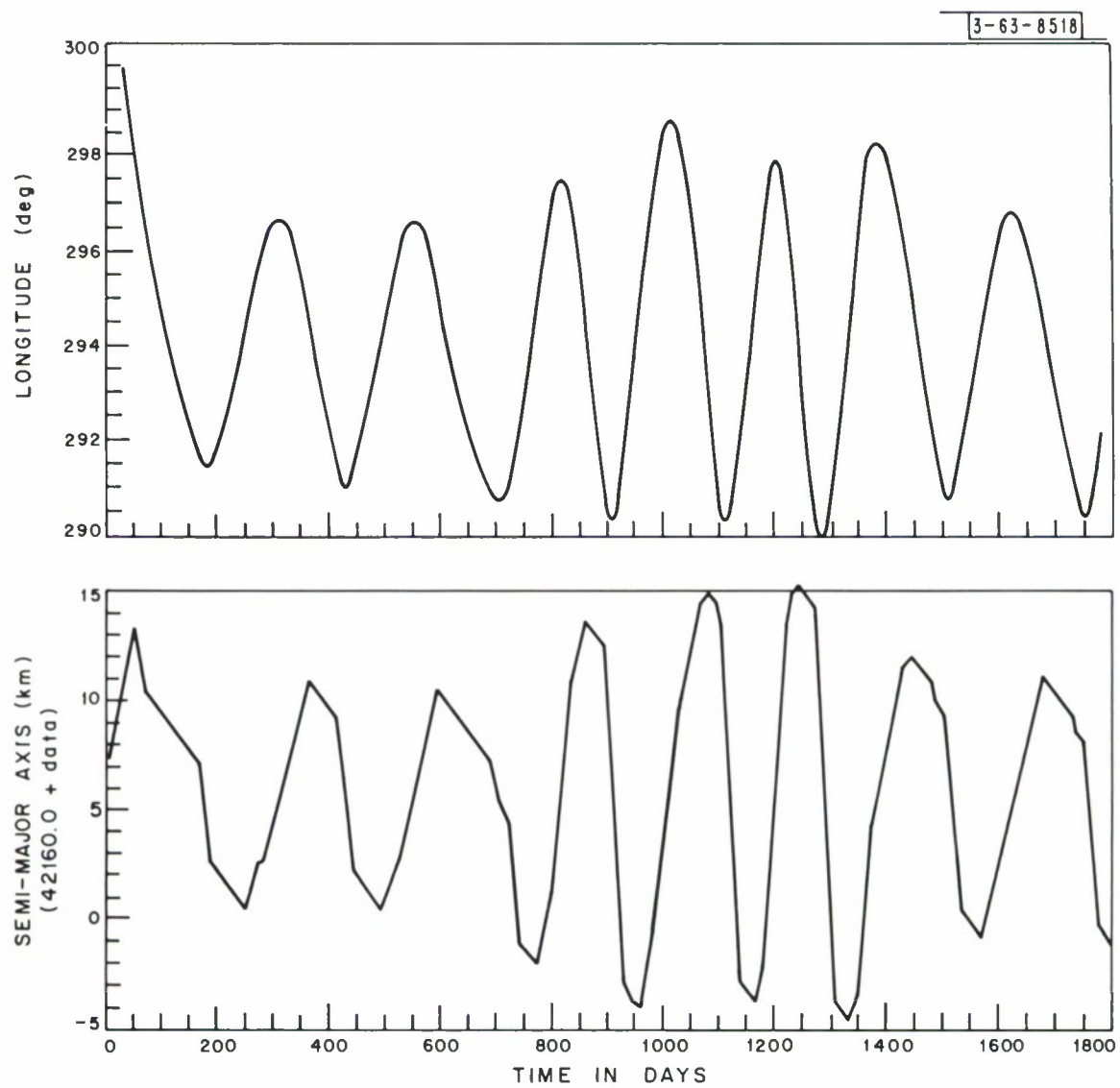


Fig. 17. Results of automatic stationkeeping using A. Braga-Illa's force scheme.

APPENDIX A
MAIN PROGRAM OF (SSS)

The main routine of the SSS program issues calls to the subroutines, solves the equations of motion, and produces the output of the program (Fig. A-1 and Appendix J). The initial section of comment cards gives a brief description of the program, its control variables and how to run it. The first executable statements are definitions of constants. The Main program calls GETINP to obtain the input data and control variables. After a sequence of housekeeping details -- moving variables to and from commons, changing variable types, and changing variables in degrees to radians -- the values necessary for the first entry into NEWFOR, FLAM, and CIND are initialized. These variables are initialized after the call to GETINP so that they are reinitialized at the beginning of subsequent sets of input data. The control variable I1 determines if the branch using second harmonics in the equations of motion or the branch using third harmonics is taken. In either case calls are issued to SSL and FLAM. SSL initializes CIND for solution of differential equations; FLAM returns the force in the λ direction. Certain logical paths determined by the input data return other variables to Main through COMMON (see Appendix D). After the call to FLAM, D2R, D2XLAM and D2THET are calculated. Both branches meet at statement 23, where three checks are made. First, is this time through the end of a day? Second, is this time through the end of ten days? Third, is this time through the end of amount of time asked for? If answer to all three questions are no, then it is end of one of the 86,400 times per day that the variables D2R, D2XLAM, and D2THET are calculated. At statement 6 the calls to FINDV, and DPNV provide T, R, THETA, XLAM, DR, DTHETA, and DXLAM for the following loop. Again a branch is taken based on I1. If it is the end of a day, the following are printed out.

TDAYS	days since beginning of this set of input data
R	radius of orbit
DR	change in radius in last loop of the day
TDEG	THETA in degrees

DTDEG	change in THETA in last loop of the day
XLDEG	XLAM in degrees
DXLDEG	change in XLAM in last loop of the day
XELONG	east longitude in degrees

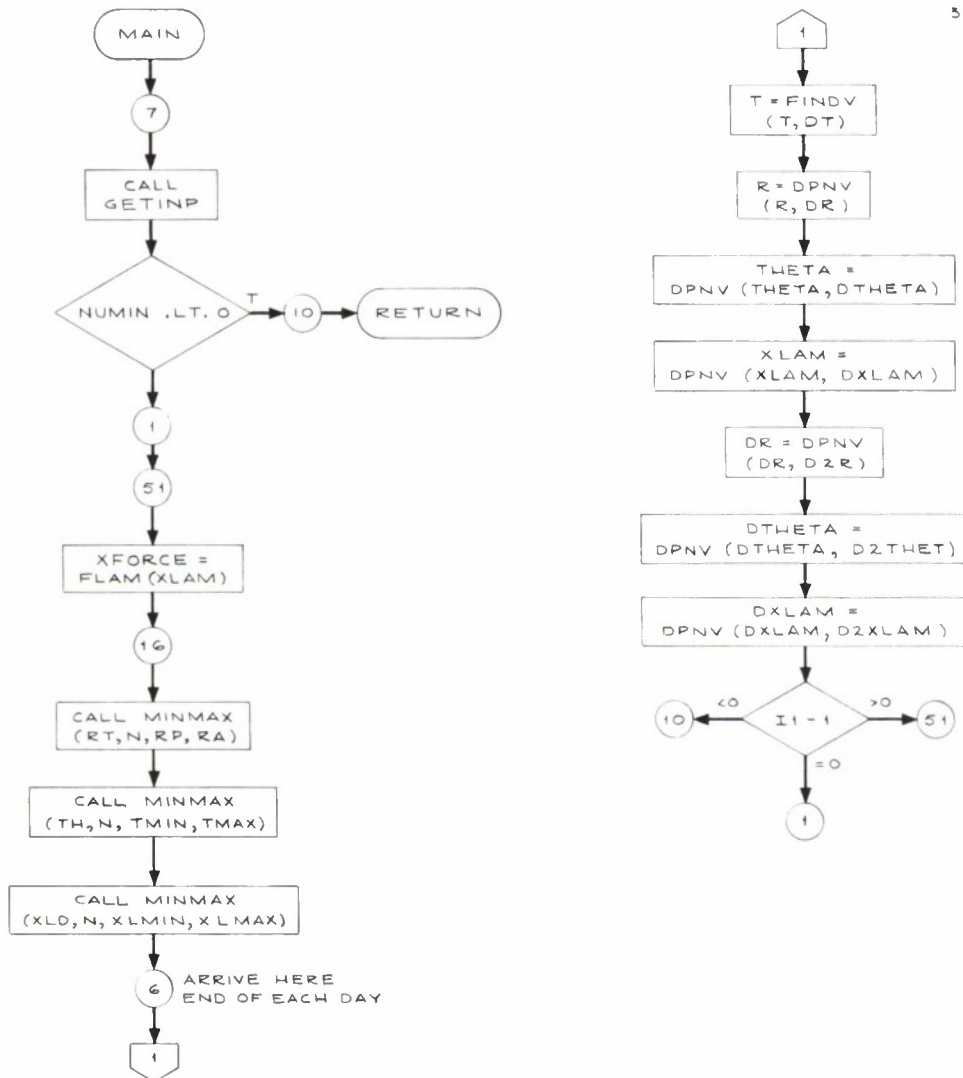
After this printout a branch is made to question three. If the answer to question two is yes also, there precedes this printout a line of print containing the following averaged over previous 10 days.

A	radius
E	eccentricity of the orbit
XI	inclination
XIM1D	minimum inclination
XLMAID	maximum inclination
DIFFER	XLMAID - XLM1D

Every ten days there is also punched a card containing TDAY, A, and XLAM, to be used by a plotting program. After this there is a branch to question 3. If the answer to question 3 is yes, and the amount of output for one set of data is finished, the program branches to 7, CALL GETINP. The last set of data contains NUM = -1, which causes program to terminate normally. Each set of data, except the last, takes 1 1/2 - 2 1/2 hours of computer time.

MAIN PROGRAM FLOW CHART

5-63-8517



APPENDIX B

The subroutine GETINP establishes the logical variables, all the input, and the modes of output. There are eight classes of integers that are used in conjunction with the logical variables: I1, I2, I3, I4, I5, I6, I7, I8. All the parameters are initialized in GETINP and then those that are to be changed are read in through a namelist setups. Using the logical parameters and the numerical parameters, the subroutine SETUP and its entry point RITOUT prepare the variable format for PRINT and/or PUNCH so that at the beginning of each output one may quickly review the particular set of parameters used. Control is then returned to the main program which then prints the titles for the output, prints and/or punches the orbital information every ten days.

I1 determines whether the calculations in the main program use second order harmonics or third order harmonics. I1 is set to 1 which selects the second order harmonics. If third is true, I2 then equals 2 and the third order harmonics are used. I2 controls the punched output. I2 is initially 1 which produces punched output. When no punched output is desired, I2 should read in as 2. The punched output is used by a separate program Z-PLOT which produces one frame of alphanumeric data and a second frame with 2 graphs longitude vs. days and semi-major axis vs. days. I3 controls the solar time correction and primitive uses of a solar sail. I3 is initialized as 1, no correction. If I3 equals 2 or linear is true, there is a linear solar time correction. If I3 equals 3 or LINRAN is true, there is a linear solar time correction with random error. If I3 equals 4 or perfect is true, the solar time is corrected perfectly. If I3 equals 5 or STEP is true, the solar time is corrected by a step-function. When I3 equals 6 or STPRAN is true, the step function correction is modified by calling the random error function. If I3 equals 7 or EPH is true, the solar time correction is done by an ephemeris sun sensor. When I3 equals 8 or EPHRAN is true, the ephemeris sun sensor correction is adjusted using the random error function. The remaining I3 values pertain to

the solar sail. When I3 equals 9 or SAIL1 is true, a fixed sail is used with a constant force along lambda. If I3 equals 10 or SAIL3 is true, a fixed sail scheme 3 is used. I4 determines which XJ22 and XLAM22 are to be used. The program is set up to use I4 equal to 2 which sets XJ22 at -1.700E-6 and XLAM at -19 degrees. There are no logical variables that go with I4. I5 establishes the force cutdown. Initially I5 equals 1, no force cutdown. When I5 equals 2, there is a force cutdown equal to the input number CUT. To obtain I5 equals 2, it can be read in directly or CUTDOWN can be read in as true. I6 and I8 are the more sophisticated uses of the solar sail. I6 equal to 1 or SAIL2 equal to true fixes the sail and allows one to also make a sun correction. I6 is initialized as 0. I6 equal to 2 or TACSAL is true the double precision function FLAM calls TACK and the tacked sail scheme is used. I8 flips the sail according to three schemes: I8 equal 1 or FLIPW equal to true fixes the sail through half the orbit and flips it off with the orbit; I8 equals 2 or FLIPAG true fixes the sail through half the orbit and flips it off against the orbit; I8 equals 3 or SAIL4 true flips the sail due to logic within the satellite. I7 determines if the orbit elements are averaged at the end of every ten days before they are punched. Initially I7 equals 2 and AVERAG is false, which loops the main program around the averaging sequence. When I7 equals 1 or AVERAG is true, the main program averages every ten days and punches the averaged figures.

The basic units for the parameters are kilometers, seconds, degrees. They are initialized as follows:

A	=	42165.5D0	radius of earth in kilometers
A1	=	0.9D0	solar absorptivity for solar sail
A2	=	0.1D0	solar absorptivity for solar sail
ALONGR	=	0.0	force along the radius to the satellite
CONST	=	-4.45D-8	kilonewtons/kilogram
			= 1.6 * (factor = 5.6)/365.25/(86.4 * 6)
			sign is negative in second and fourth coordinates

CUT	=	0.0	
DATE	=	0.0	
E	=	0.0	
FORCCT	=	0.0	force along theta
FORLAM	=	1.0	force along lambda
IRUN	=	0	a counter so one can keep track of the various computer runs
ITHDAY	=	10	frequency in days of print and/or punch execution
NUM	=	6	number of impulses given (used in FLAM)
OFFSET	=	-90.0D0	angle by which sail can be offset in the TACK scheme
OL	=	0.0D0	argument of perigee for first orbit, in degrees
TEND	=	1800.0D0	number of solar days program is to run for
XDISP	=	70.0D0	specific impulse of thruster in seconds
XI	=	0.0D0	inclination of the orbit in degrees
XLAMD	=	300.0D0	degrees of longitude measure east of Greenwich
XLIMD	=	297.0D0	minimum angle for thrusting, simple scheme
XLIMXD	=	350.0D0	maximum angle for thrusting, simple scheme
XMASS	=	132.0D0	mass of satellite in kilograms

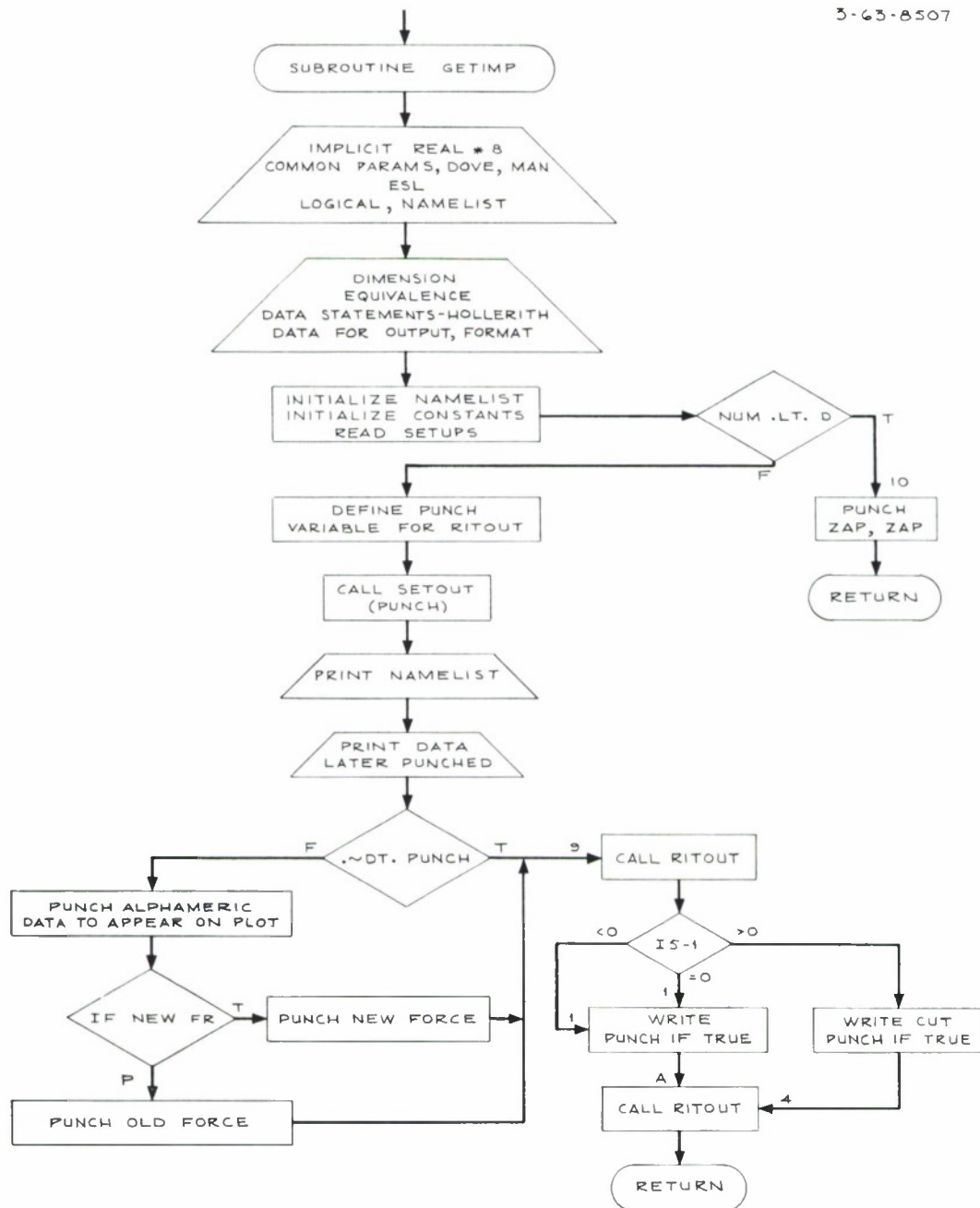
In addition there is a logical parameter, NEWFR, that has no numerical equivalent. NEWFR is initially FALSE. When NEWFR is true, the SUBROUTINE NEWFOR is called by FLAM and the magnitude of the impulse is modified accordingly.

To change any of the parameters or logical variables the NAMELIST/SETUPS/ is used &SETUPS is on the first card beginning of column 2. The variables desired changes are then punched separated by commas. To set a logical variable to TRUE just use T for example: NEWFR = T would be

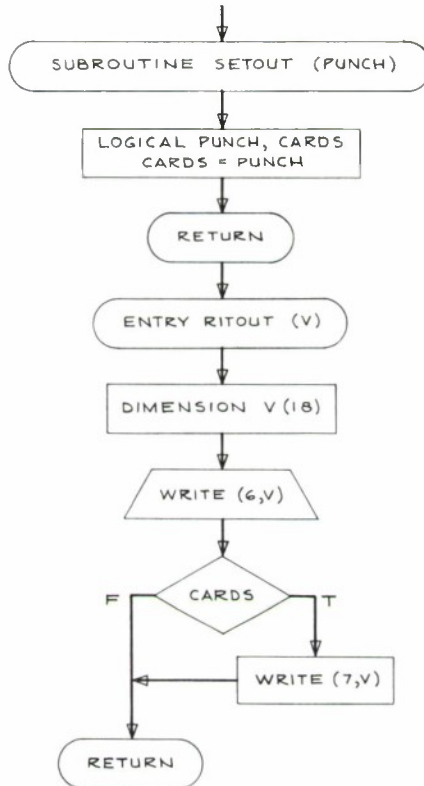
sufficient to set NEWFR true. Do not punch in column 73 or beyond. When using a second card begin in column 3. Do not follow the last variable with a comma; instead leave one or more spaces and punch &END. The set of cards with the SETUPS changes is then placed behind the //FT05F001 control card.

The parameters are printed and/or punched according to a variable FORMAT routine, RITOUT. RITOUT is an ENTRY point in SETOUT. The first call to SETOUT establishes whether or not to punch when RITOUT is called. RITOUT is called for each I variable. To add an I variable to the program, IY, where Y is an integer greater than 8, eight additions to GETINP are necessary. First a comment card to describe the IY variable and the X uses to which IY can be put. Secondly, it must be initialized. The remaining additions are to setup the call to RITOUT. There are X + 1 dimension statements as follows: Dimension VFIY(16,6) and X statement like dimension XVFIYX(16). For each X there is also an EQUIVALENCE statement such as: EQUIVALENCE (XVFIY(1), VFIY(1,Y)). For each value IX can assume there must be a data statement in Hollorith form containing the alphanumeric data you desired printed and/or punched to describe the alternative used. You need to add IX to the 101 format statement, and one call to RITOUT, for example: Call RITOUT (VFIY(1,X)).

GETINP thus establishes the particular parameters and logical choices for one computer run, which takes approximately two hours on the IBM/360 with TEND = 1800.



3-63-8508



APPENDIX C

INTEGRATION FUNCTION AND SUBROUTINE

The double precision function CIND(K, A, B) and its related functions FINDV, DPNV, and SUBROUTINE SSL find the first and second derivatives and calculate new values for R, THETA, XLAM, DR, DTHETA, DXLAM, and T for each cycle through the main program. The call to SSL sets up the constants for SSL. FIND calls CIND with K = 0 which routes the control to the proper place to find a new T. For the other variables DPNV is used. In the computer GO TO statements only the first two solutions are used by this program. The sequence of four subprograms was written in another language outside Lincoln, so there are more choices than this program uses.

The integration formula used is the Adams four-point formula:

$$y_{n+1} = y_n + \frac{h}{24} (55y'_n - 59y'_{n-1} + 37y'_{n-2} - 9y'_{n-3}) \quad (C-1)$$

where $h = x_{n+1} - x_n$ is the constant incremental change of the independent variable x , y being a function of x .

Starting formulas are needed to obtain the first three points. The following set of formulas is used to obtain y_1 , y_2 , and y_3 :

$$y_1^{(1)} = y_0 + hy_0' \quad (C-2)$$

$$y_1^{(2)} = y_0 + \frac{h}{2} (y_1^{(1)'} + y_0') \quad (C-3)$$

$$y_1^{(3)} = y_0 + \frac{h}{6} (2y_1^{(2)'} + y_1^{(1)'} + 3y_0') \quad (C-4)$$

$$y_2^{(1)} = y_1^{(3)} + \frac{h}{2} (3y_1^{(3)'} - y_0') \quad (C-5)$$

$$y_2^{(2)} = y_1^{(3)} + \frac{h}{12} (5y_2^{(1)'} + 8y_1^{(3)'} - y_0') \quad (C-6)$$

$$y_3^{(1)} = y_2^{(2)} + \frac{h}{12} (23y_2^{(2)'} - 16y_1^{(3)'} + 5y_0') \quad (C-7)$$

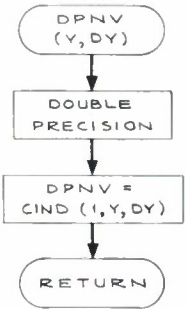
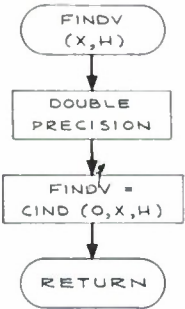
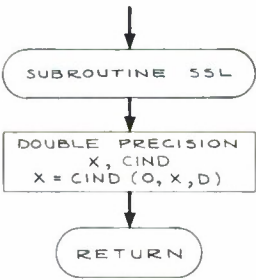
$$y_3^{(2)} = y_2^{(2)} + \frac{h}{24} (9y_3^{(1)'} + 19y_2^{(2)'} - 5y_1^{(3)'} + y_0') \quad (C-8)$$

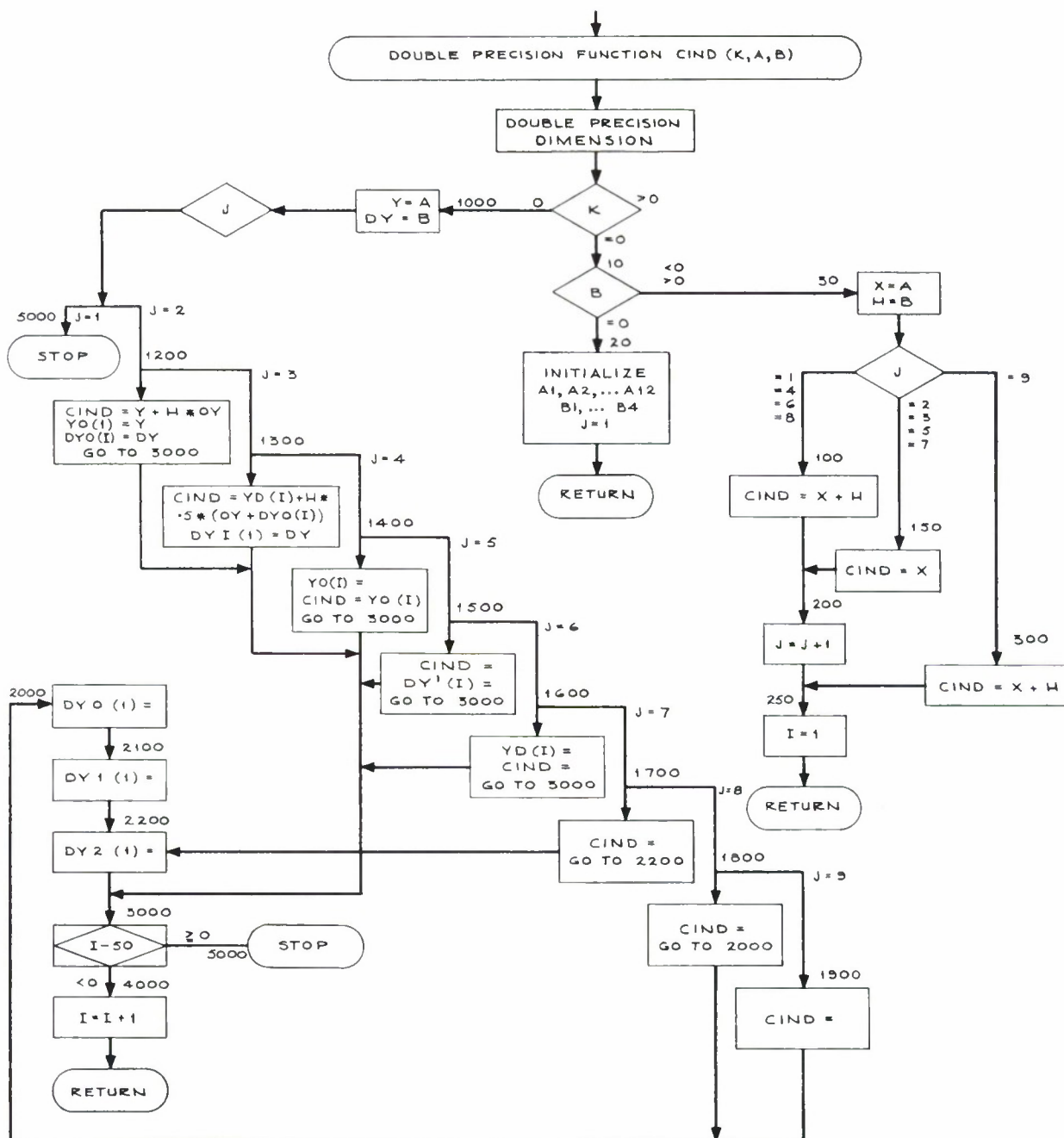
where the superscripts indicate the order of the approximation. The past values of the derivatives are stored in reserved blocks. For a more complete discussion of the starting equations, the reader is invited to contact the author. In particular, a substitution of equations (1) and (2) and similar expressions for their derivatives into equation (3) will show that equation (3) reduces to the first four terms of the Taylor series.

The truncation error of formula I is $\frac{251}{720} h^5 y^{(5)}(s)$, where $x_n < s < x_{n+1}$.

The accumulation of the increments $(\Delta y)_n = \frac{h}{24} (55y_n' - 59y_{n-1}' + 37y_{n-2}' - 9y_{n-3}')$ is performed in double precision to substantially reduce the round-off error.

Complete underflow-overflow testing is made. An underflow is corrected and a zero stored.





APPENDIX D

DOUBLE PRECISION FUNCTION FLAM (XLAM)

The double precision function FLAM(XLAM) computes a force, FLAM, simulating the force of the thruster or solar sail depending on the logic with which the program is run. FLAM is returned to the main program. The force is applied NUM times an orbit, controlled by the gate K. When K is positive the gate is closed returning both force and FLAM as 0. When K is negative control passes to label 8. The various routes for the logic depend on the force scheme used for the particular run in question. The variable 15 controls the cutdown; when 15 is 1 there is no cutdown. When 15 is 2, the force is cut by the factor CUT, as long as LLL is also positive. LLL is the control variable whose sign is determined by the satellites position within a specified band, delineated by XLIM and XLIMX. LLL is negative when XLAM is inside the band and positive when XLAM is outside the band. 13, for values 1-8, controls the type of clock error returned by CORR(TEL). From this an XTEMP is calculated. XLAM is the actual position of the satellite; XTEMP is where the satellite's logic thinks it is. The old force calculations use LLL, whereas the new force calculations use the subroutine NEWFOR. As long as the counter M is less than NUM, control returns to the main program with FLAM = FORCE. Once an orbit the fuel consumption, the total number of impulses, JFK, CORR the corrected longitude in degrees are printed out. The fuel consumption is kept as a running sum. After the fuel report is printed, on the last loop during which there is thrusting, K is set positive and the other counters are initialized for the next orbit.

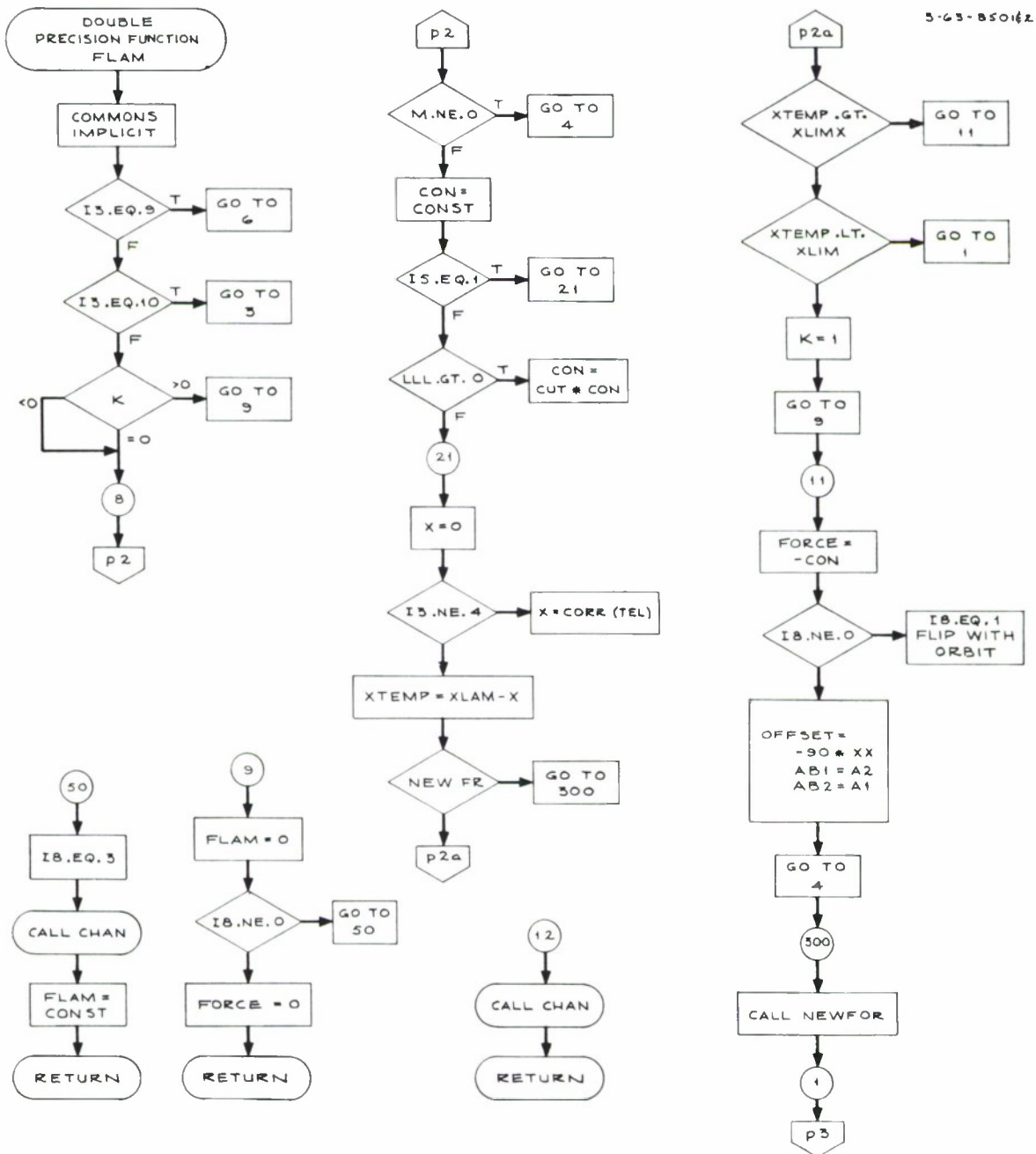
For the various schemes of solar sailing, there are deviations from the above possibilities. When K is positive and the sail is flipped off, FORCE = CONST and CHAN is called. When a simple sail is used FLAM = CONST and return. For this sail there is a constant force along λ . When a fixed sail is used, 13 = 10, FLAM = CONST and TACK or CHAN is called depending on 16.

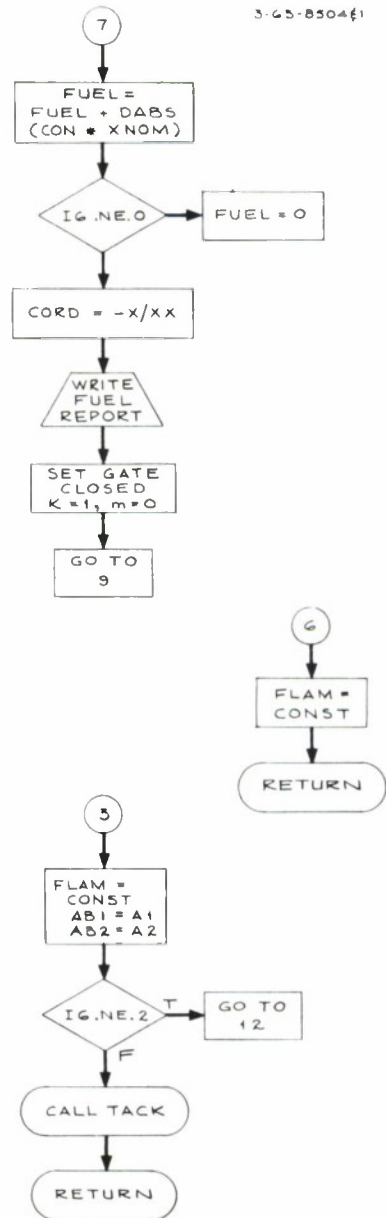
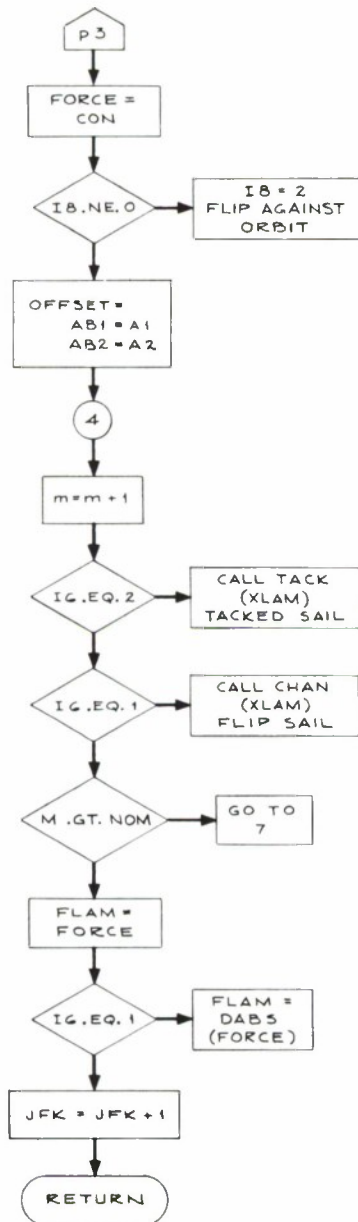
Tack calculates the forces in other directions as in addition to the force along λ the FLAM returns. Control returns to the Main program.

When K is negative (or zero) the force FLAM is calculated including the CUTDOWN or not. The first time through $CON = CONST$, XLAM is correlated by $CORR(TEL)$. It is then determined if the modified XLAM lies within the band, if so $FLAM = 0$ and control is returned to the main program. If the true longitude is outside the desired stationkeeping range control passes to statement 1 or 11 within FLAM. If the satellite is too far east, $FORCE = -CON$, the dial can be flipped with the orbit if that scheme is used, or the sail can be offset -90° if that scheme is used. If the satellite is too far west, $FORCE = CON$, the sail can be flipped against the orbit, $OFFSET$ can be 0° . In either case control passes to the same point 4, where K will return control for the rest of the orbit as long as the impulses are being given.

At statement 4 M dutifully counts the impulses. CHAN or TACK or neither is called according to 16. If M exceeds NUM, the fuel consumption report is issued. On all the sail schemes the fuel is set to zero. And the counters are reset for the next orbit as in the non-sailing schemes.

The double precision function $FLAM(XLAM)$ returns to the main program NUM times an orbit the force along λ necessary to stationkeep within a specified range according to various force schemes using thrusters and solar sails.





APPENDIX E

SUBROUTINE NEWFOR

Subroutine NEWFOR is an alternative method of calculating the force necessary to keep a satellite in a synchronous orbit using the bilateral thrusting method proposed by A. Braga-Illa.

A. Theory

Figure 18 shows a sample of longitude vs. time, and the regions $\pm A$, $\pm B$, $\pm C$. The value of the force, con , at any point will be determined by which region we are in at that point and where we were before.

- I In region C and $-C$, $\text{con} = C_3$ and $-C_3$ resp.
 In region B and $-B$, $\text{con} = C_2$ and $-C_2$ resp.

For regions A and $-A$ there is a complicated scheme for determining the value:

- II If we get into region $+A$ via $+B$ or $-A$ via $-B$, then $\text{con} = 0$.
 If $+A$ via $-A$, $\text{con} = C_1$ or $-C_1$ resp.

III As soon as we get the sequence $(A-A+A)$ or $(-A+A-A)$ we start damping the force basically as follows. Assume we get the sequence $+A-A+A$ and then the curve never leaves the $+A$, $-A$ region.

Region:	A	B	$+A$	$-A$	$+\bar{A}$	$-A$	$+A$	$-A$	$+A$	$-A$	$+A$	$-A$
Force:	C_1	C_2	0	$-C_1$	$\frac{C_1}{2}$	$-\frac{C_1}{2}$	$-\frac{C_1}{4}$	$-\frac{C_1}{4}$	$\frac{C_1}{8}$	$-\frac{C_1}{8}$	$\frac{C_1}{16}$	$-\frac{C_1}{16}$
Region:	$+A$	$-A$	$+A$	$-A$	$+A$	$-A$						
Force:	$\frac{C_1}{32}$	$-\frac{C_1}{32}$	0	$-\frac{C_1}{32}$	0	$-\frac{C_1}{32}$						

where \bar{A} is the region where we start damping. After (*) we do not damp but keep $\pm C_1/32, 0$ as long as we stay in $+A-A$ and do not stray into B.

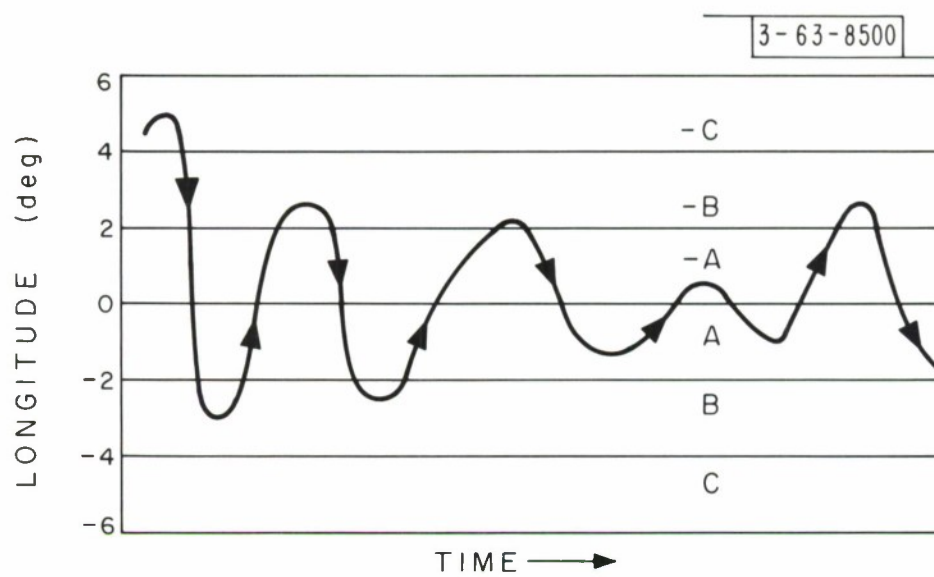


Fig. 18. Longitude vs time.

The most practical way to show the deviations from the basic scheme is via figure 19. Note the following:

(1) When we get into region C (see arrow 1) we start over in the sense that no damping occurs until we get the $\pm A \mp A \pm A$ sequence again (see arrow 2).

(2) When damping has started and we go into region $\pm B$, and then back into $\pm A$ we keep the damped value that we had before going into region $\pm B$, with the appropriate sign (see arrows 3, 4).

(3) When we reach the maximum damping stage ($\pm C_1/32, 0$) and then go into region $\pm B$, upon re-entry into $\pm A$ the force = 0 by part II and if we then go into $\mp A$ from there, force = $C_1/32$ with the appropriate sign (see arrows 5, 6).

B. Data Cards

To specify that subroutine NEWFOR is to be used to calculate the force we must specify four quantities on the data cards:

$$\text{NEWFR} = T, C_1 = \quad, C_2 = \quad, C_3 =$$

where C_2, C_3 are the values assigned to the force in regions B and C respectively and C_1 is the non-zero undamped value of the force in region A.

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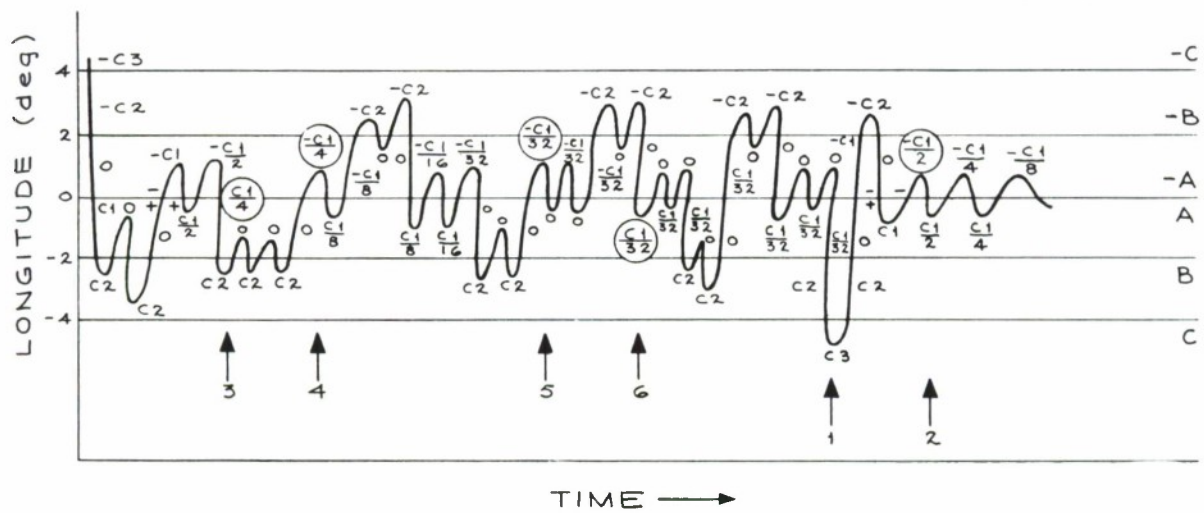
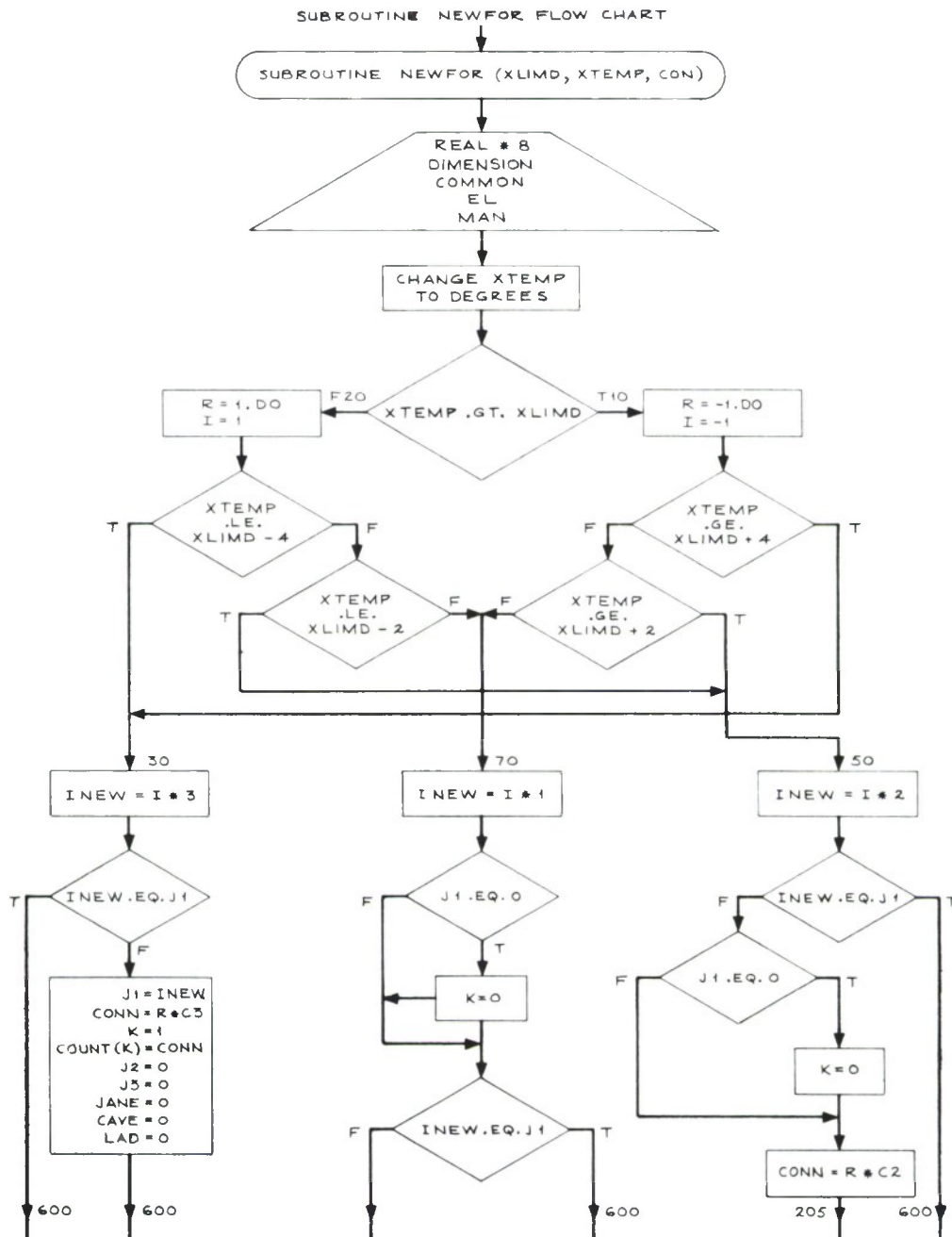
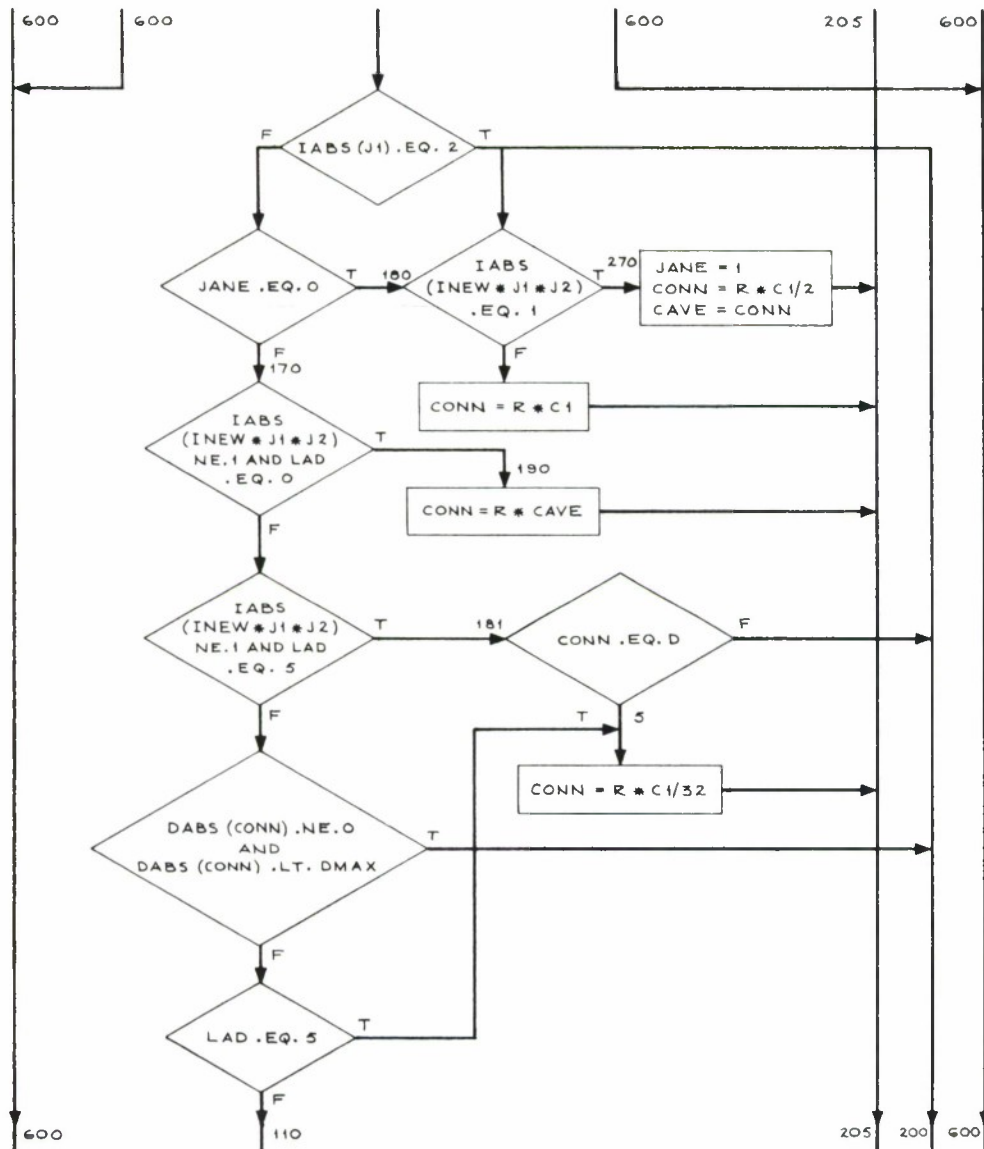
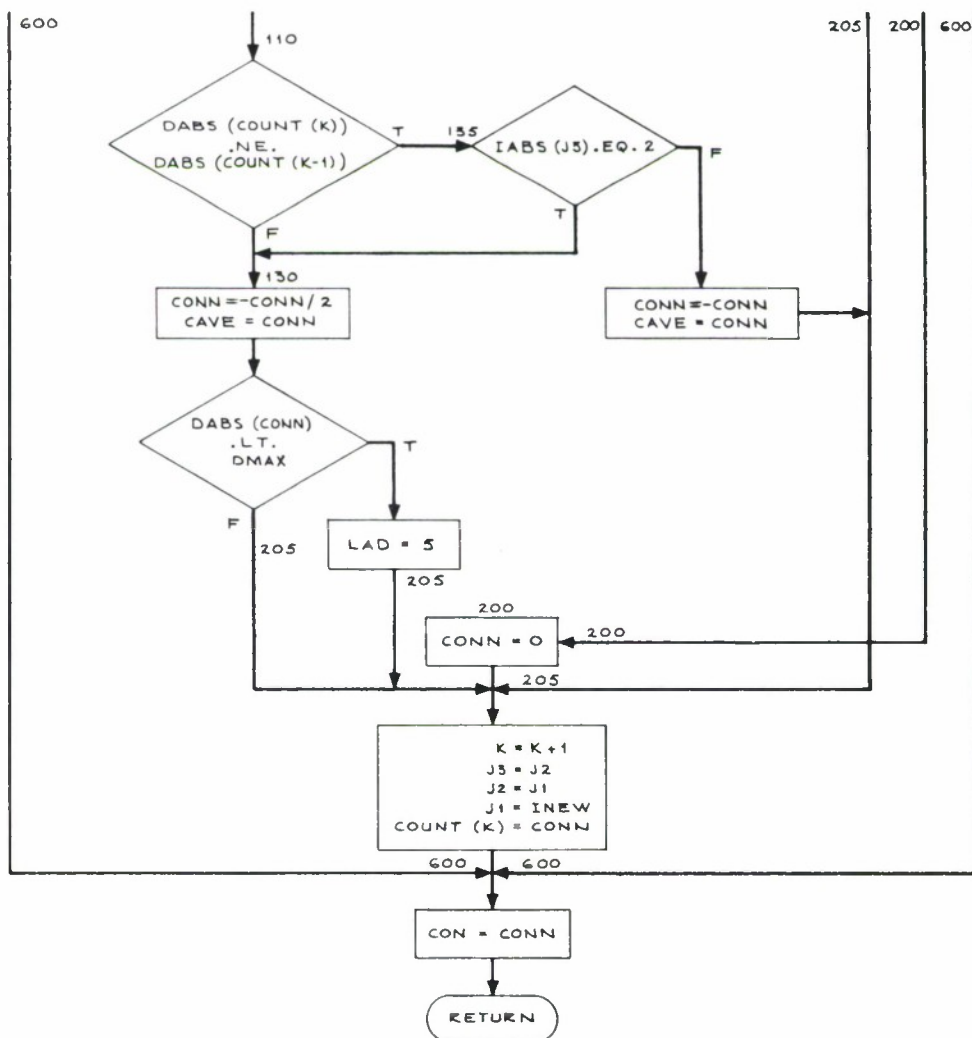


Fig. 19. Example of possible use of force scheme of A. Braga-Illa.







APPENDIX F

DOUBLE PRECISION FUNCTION CORR (TEL) returns to FLAM an angle by which the longitude XLAM is thereby modified. In FLAM: $X = \text{CORR}(\text{TEL})$, $\text{XTEMP} = \text{XLAM} - X$. The program calculates the apparent right ascension of the sun and the time of ephemeris transit, good for the year 1967, and accurate to about four seconds of arc of the right ascension and about 1.5 seconds on the ephemeris transit. There are eight options in this program determined by the I3 variable. Two of the options call the SUBROUTINE CURT, and two of the options call the SUBROUTINE ALVISE. When $I3 = 4$, the logic in FLAM bypasses the call to CORR.

If I3 is 7 or 8, immediately after converting TEL from seconds of days, the program branches to 110. Day is modulated by 365, and CURT is called to determine VALUE. If $I3 = 8$, VALUE is changed to radians and a random error is added on. If $I3 = 7$, CORR is simply VALUE in radians and control is returned to FLAM. CURT picks up an ephemeris sun sensor value from the step function tables stored in the BLOCK DATA/SPETS/. These tables simulate the logic in an electronic sun sensor within the satellite.

For the other options, the program calculates the mean anomaly, sun's coordinates, referenced to a geocentric coordinate system, right ascension of the sun, right ascension of the mean sun, and DIFF the difference between the right ascension of the sun and the right ascension of the mean sun added to 6.28 when $|\text{DIFF}|$ is less than 3. Now the program branches according to I3 to calculate VALUE. In each case after VALUE is calculated, control goes to statement 31 where $\text{CORR} = \text{DIFF} - \text{VALUE}$.

If I3 is 5 or 6, VALUE is returned from a call to SUBROUTINE ALVISE. When $I3 = 6$ DIFF has a random error added in. ALVISE as a subroutine operates identical to CURT, looking up a value in a BLOCK DATA. When control is returned to CORR, CORR branches to 31.

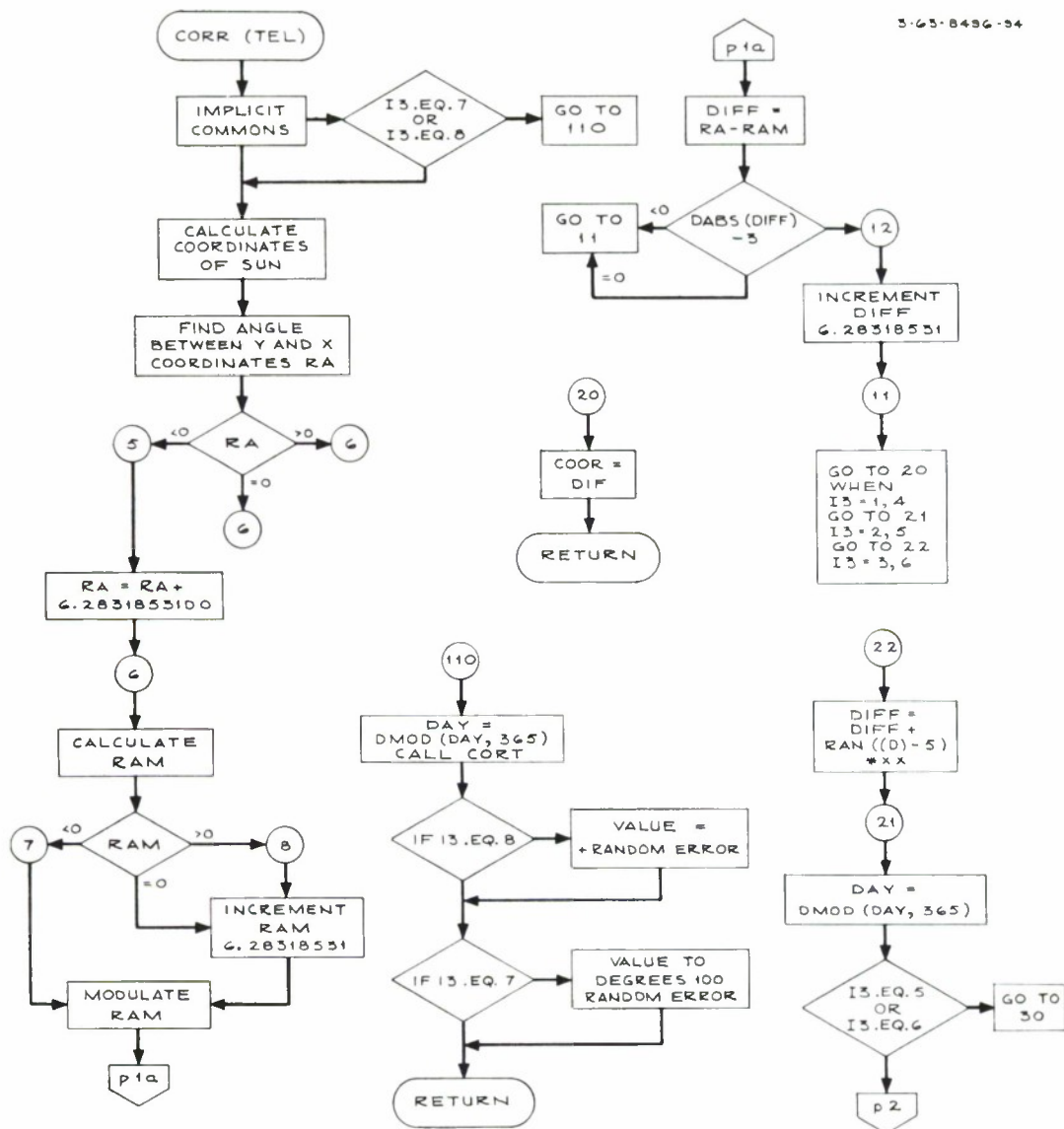
If $I3 = 1$, $\text{CORR} = \text{DIFF}$ and control is returned to FLAM. When $I3 = 2$, DAY is modulated by 365. The formula for VALUE is determined by which part of the year DAY falls.

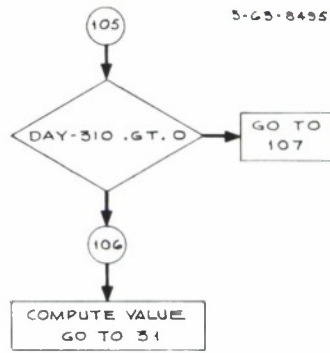
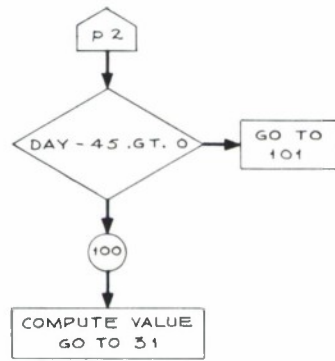
<u>Day</u>	<u>Value</u>
0 - 45	100 VALUE = .08D0 * (DAY + 2.25D0) * XX
46 - 129	102 VALUE = -.0584D0 * (DAY - 111.D0) * XX
130 - 211	104 VALUE = .033D0 * (DAY - 162.D0) * XX
212 - 310	106 VALUE = -0.0584D0 * (DAY - 240.D0) * XX
310 - 360	107 VALUE = .08D0 * (DAY - 363.D0) * XX

After each calculation of VALUE control branches to 31. When I3 is 3, DIFF is modified using the random error function, then logic follows the same for I3 = 2. There is a loop for I3 = 4, although CORR is not called when I3 = 4, to avoid an error message from the compiler.

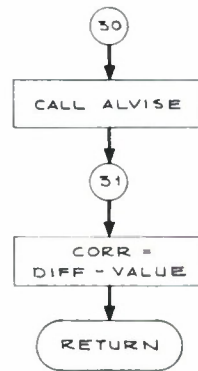
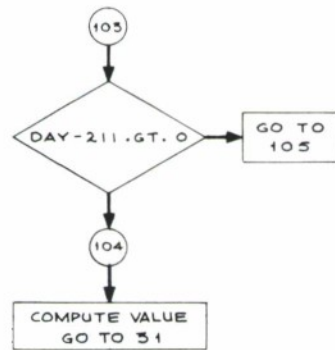
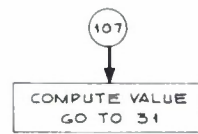
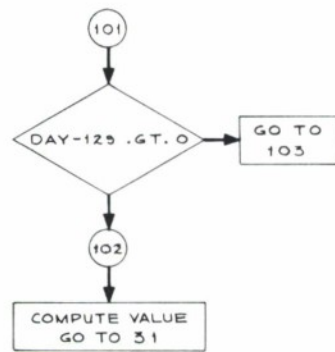
FUNCTION CORR FLOW CHART

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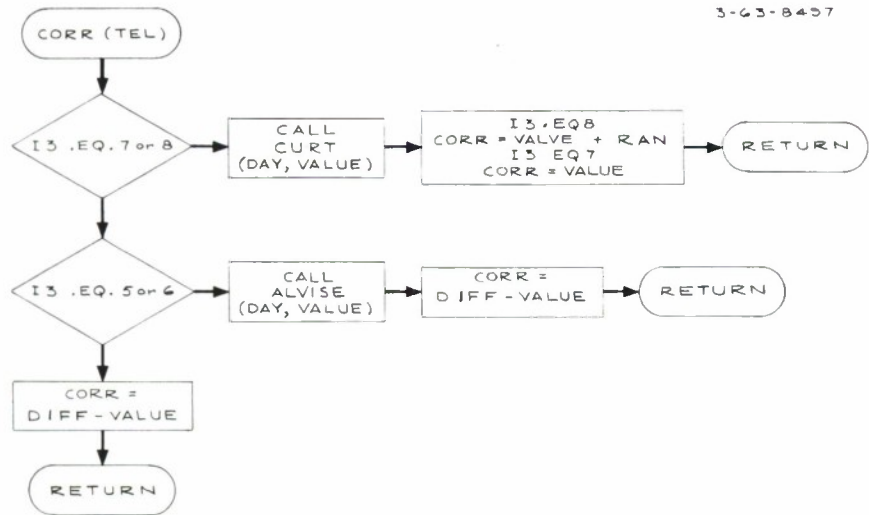




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3-63-8457



APPENDIX G

The SUBROUTINE TACK (XLAM) and its ENTRY CHAN (XLAM) return to MAIN via FLAM three forces, FORCCT, FORLAM, and ALONGR, which are used to calculate D2THET, D2XLAM, and D2R respectively. The ENTRY CHAN(XLAM) fixes the sail along the radius to the satellite from the earth. According to the 18 variables, the sail remains upright on the radius or is flipped throughout part of the orbit. FORCCT is the solar force exerted by the sail in the direction of theta. FORLAM is the solar force exerted by the sail in the λ direction. ALONGR is the solar radiation force exerted along the radius. L. Black has pointed out that the most efficient sail would be one of high thermal emissivity on both sides with a high solar absorptivity on one side and a high solar reflectivity on the other. A sail made out of aluminized microsheet stuck on an aluminum sheet painted black on the back side would be a good approximation to this. The average force along the orbit for a sail whose plane is coincident with the orbit radius vector and the north-south direction is given approximately by

$$F = \frac{AS}{2C}$$

where

- A = is the area of the sail
- S = solar constant
- C = velocity of light

The sail placed on angle bisector of angle formed by sun and orbital tangent in direction of the motion is the most efficient sail; the program allows also an offset from this position, OFFSET.

When the solar rays hit the sail they generate two forces which must be resolved into the three components needed. The two forces are one a force perpendicular to the sail itself and opposite to the side of the sail hit by the solar radiation; the second is a force along the sail due to its various

absorptivity properties. Both sides of the sail are reflective when the sail is "tacked" and when it is flipped. When ENTRY CHAN(XLAM) is used, the sail is fixed on the radius from the earth to the satellite, and one side is highly reflective while the other side is highly absorptive. The three components of the generated forces are the force along the orbit, FORLAM; the force along the radius vector of the satellite, ALONGR, and the force along λ , FORLAM.

In this subroutine all the vectors are unit vectors and lie in the orbital plane of the satellite. The coordinate system is geocentric, with the X-axis lying along Aries. XNORMA, YNORMA are the coordinates of the normal to the sail itself. The coordinates of the sun are calculated as in Moulton pp 171, and then the vector is projected into the orbital plane with coordinates XPROJ, YPROJ. XFORCE, YFORCE are the coordinates of the orbital tangent. XSAIL, YSAIL are the coordinates of the normal to the sail. XRESAL, YRESAL are the coordinates of the sail itself. The angle between the sail and the radius is ANCHOR. DOTPRO is the cosine of the angle between the normal to the sail and the projection of the sun. The absolute value of the cosines is used to simplify the logic in determining quadrants. ROCK is the absolute value of the cosine of the angle between the sail and the radius. TRAP is the absolute value of the cosine of the angle between the sun and the normal to the sail. When DOTPRO is less than zero the sail is receiving the solar force on its absorptive side. POT is the force exerted perpendicular to the sail due to the solar radiation pressure. $POT = (2 - A) * DOTPRO ** 2$. ALOSAL is the force along the sail due to a non-perfect sail.

$$ALOSAL = -A * TRAP * (XSUN * SRESAL + YSUN * YRESAL)$$

There is a logical sequence to test whether the sail is "ON" with the flipped sail logic. If it is on one continues, if not the program branches to 40, where a sequence of equations set up the small forces when the sail is turned "OFF". Four variables are used to set up the signs of the forces according to the quadrant the sail, sun, and orbital tangent are in. They are COD, FISH, HERRING, and SALT. They are initially +1 and are made -1 when necessary. The solar

force is exerted opposite to the normal to the sail when the angle between the sun's projection and the sail is between zero and 180; the solar force is exerted along the normal to the sail when the angle between the sun's projection and the sail is between 180° and 360°. The coordinates of the solar force are XSORAF, YSORAF. HERRING is negative when the angle between the solar force and the orbital tangent is greater than 180. SALT is negative when the angle between the sail and the orbital tangent is greater than 180. COD is negative when the angle between the solar force and the radius is greater than 180. FISH is negative when the angle between the radius and the sail is greater than 180. The three components of the solar radiation force acting on the satellite are:

$$\left. \begin{aligned} \text{FORLAM} &= \text{HERRING} * \text{POT} * \text{ROCK} + \text{SALT} * \text{ALOSAL} * \text{BOAT} \\ \text{ALONGR} &= \text{COD} * \text{POT} * \text{BOAT} + \text{FISH} * \text{ALOSAL} * \text{ROCK} \\ \text{FORCCT} &= -\text{AB1} * \text{TRAP} * \text{ZSUN} \end{aligned} \right\} \begin{array}{l} \text{See} \\ \text{Fig-} \\ \text{ure} \\ 20 \end{array}$$

When the sail is flipped to its "OFF" position there are slight forces acting upon the satellite. DOTPR1 is the cosine of the angle between the sun (not its projection) and the sail. DOTPR2 is the absolute value of ZSUN. FORCCT is negative when ZSUN is positive. Otherwise the three forces are:

$$\begin{aligned} \text{FORLAM} &= \text{AB1} * \text{DOTPR2} * \text{DOTPRO} \\ \text{ALONGR} &= -\text{AB1} * \text{DOTPR2} * \text{DOTPR1} \\ \text{FORCCT} &= (2 - \text{AB1}) * \text{DOTPR2} \end{aligned}$$

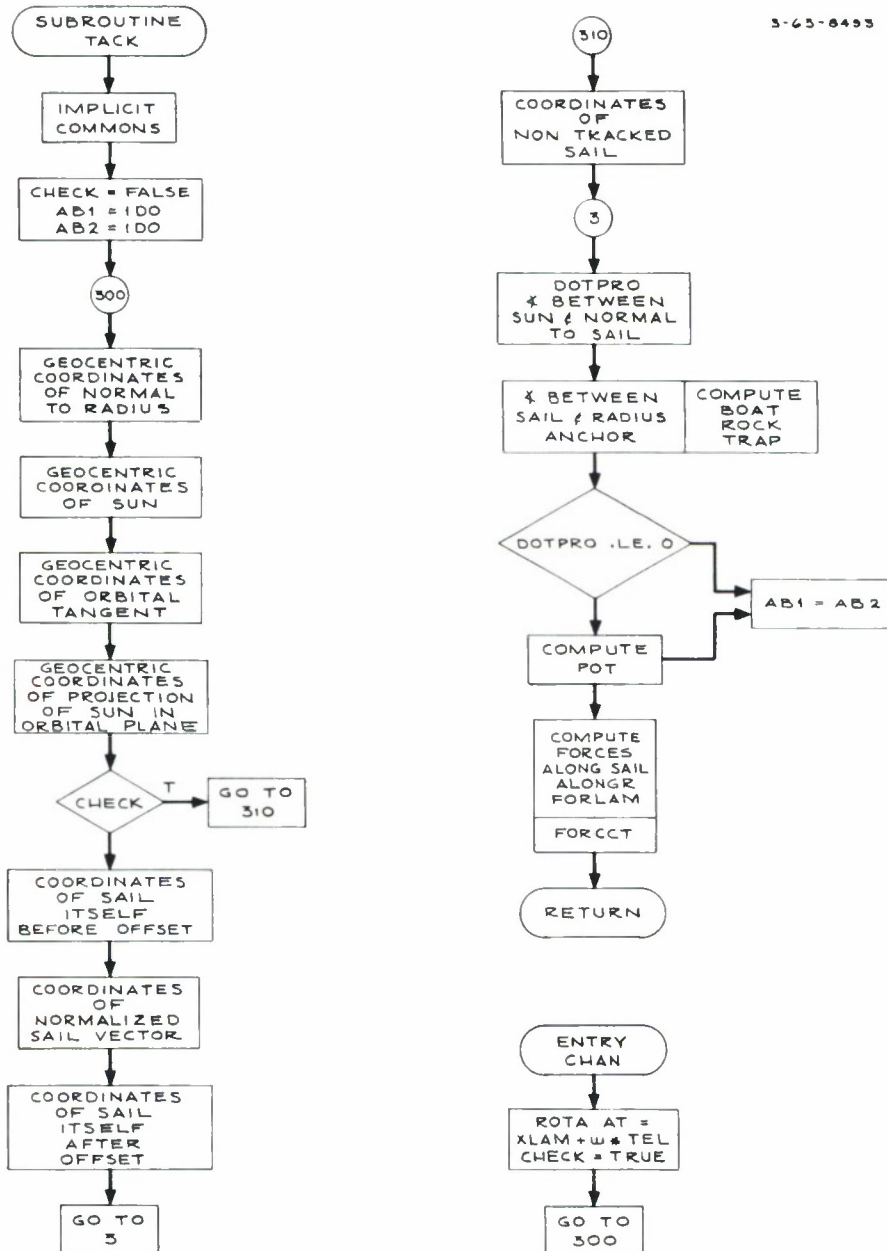
AB1 is the absorptivity received by the program as A1 and A2 and set equal to the proper one depending on the quadrant one is in. Following is a diagram of the forces for an offset sail in one quadrant.



Fig. 20. VECTOR DIAGRAM OF FORCES acting on solar sail.

SUBROUTINE TALK FLOW CHART

3-63-8493

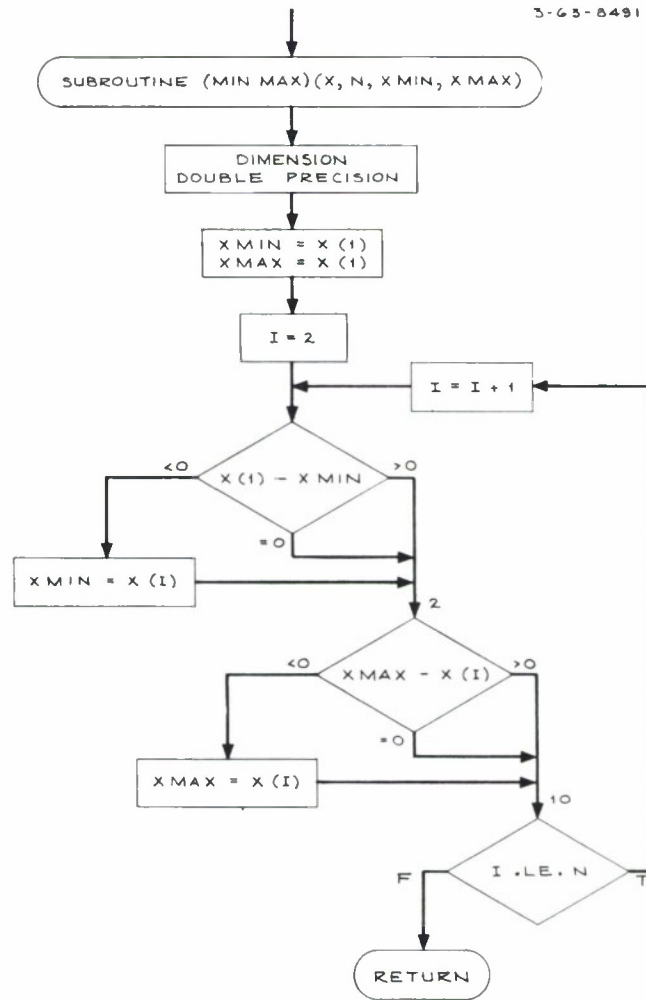


APPENDIX H

SUBROUTINE MINMAX (X, N, XMIN, XMAX)

Subroutine minmax is designed to find the maximum and minimum value of a set of N numbers.

3-63-8491

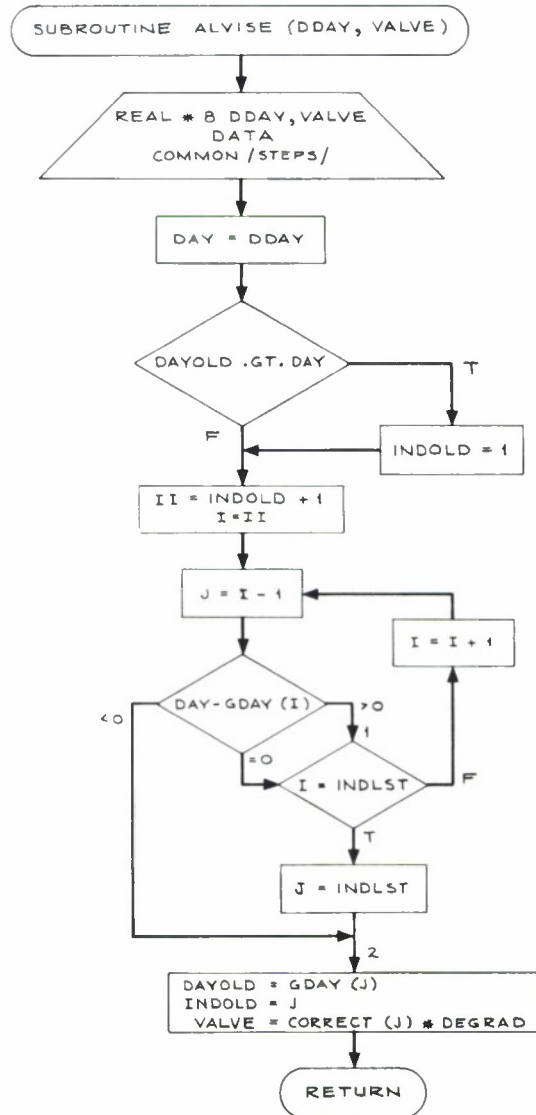


APPENDIX I

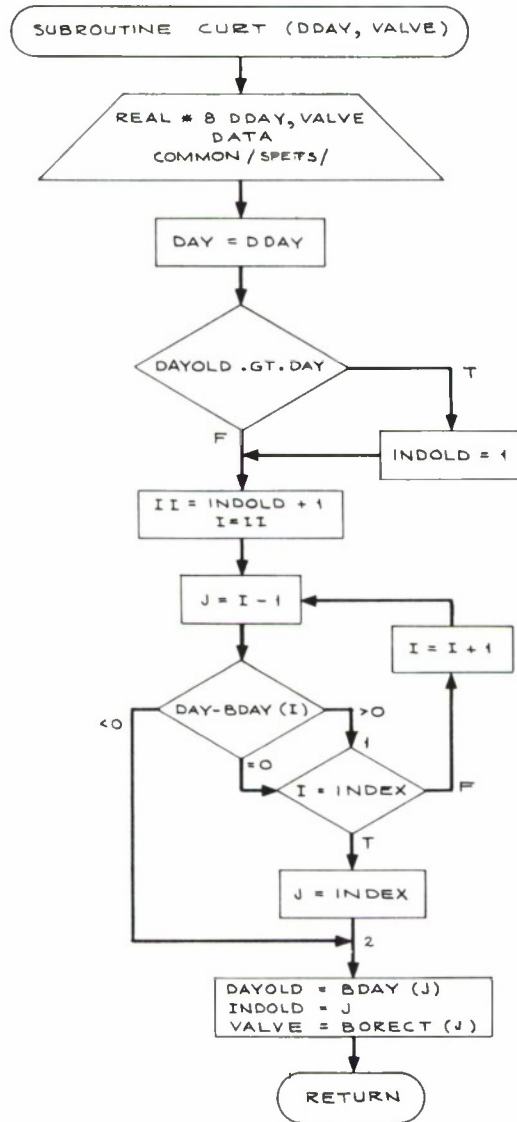
SUBROUTINES ALVISE AND CURT

SUBROUTINE ALVISE contains the information for the step function approximation to the equation of time used for the LES-6 satellite.

SUBROUTINE CURT contains the information for the step function approximation to the equation of time used for the LES-6 satellite analemma sensor.



3-43-8480



APPENDIX J

LISTING OF THE PROGRAM

LEVEL 2 FEB 67

OS/360 FORTRAN H

DATE 68.067/17.15.43

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,EBCDIC,XLIST,XJOECK,LDA3,MAP,NDEOIT,JD

```

C THIS VERSION OF Z800 HAS NO REAL I6 OPTION IN IT.
C XLAM MEASURED EAST OF GREENWICH
C
C IF I1 = 1, 2ND ORDER HARMONICS USED
C IF I1 = 2, 3RD ORDER HARMONICS USED
C IF I2 = 1, PUNCH OUTPUT
C IF I2 = 2, NO PUNCH OUTPUT
C I3 CONTROLS SUN CORRECTION
C
C I3 = 1 NO CORRECTION
C I3 = 2 CORRECTED LINEARLY
C I3 = 3 CORRECTED LINEARLY W/ RANDO
C I3 = 4 CORRECTED PERFECTLY
C I3 = 5 CORRECT BY STEP FUNCTION
C I3 = 6 CORRECTED BY S-F W/ RANDOM
C I3 = 7 CORRECT BY EPHEMERIS SUN SENS
C I3 = 8 CORRECT BY EPHEMERIS SS W/R
C I3 = 9 1ST SUN SAIL CORRECTION
C CONSTANT FORCE ALONG LAMDA
C I3 = 10 SOLAR SAIL3 FIXED SAIL
C
C IF I4 = 1 XJ22=-1.816E-6 AND XLAM22 = -15.0*-01745
C IF I4 = 2 XJ22=-1.700E-6 AND XLAM22 = -19.0*-01745
C IF I5 = 1, NO FORCE CUTOOWN
C IF I5 = 2, FORCE CUT BY INPUT NO. CUT
C I6 = 1 SOLAR SAIL2
C IF I7 = 1, PLOTTING AVERAGE OF XLAM FOR 10TH DAY
C IF I7 = 2, PLOTTING XLAM AT END OF EVERY 10TH DAY
C IF I8 = 3, SAIL IS FLIPPED BY LOGIC IN SATELLITE
C
C IF NEWER=TRUE, USE NEW FORCE CALCULATIONS
C IF NEWER=FALSE, USE OLD FORCE CALCULATIONS
C
C CONST IS CALCULATED IN THE FOLLOWING MANNER--CONST=F/M EXPRESSED IN
C KILOGNEUTONS PER KILOGRAM. FOR OUR CASE
C  $1.6 * (FACTOR = 5.6 / 365.25 / (86.4 * 6)) = 4.45E-8$ 
C THE SIGN SHOULD BE NEGATIVE IN THE SECOND AND FOURTH COORDINATES
C
C Z800 IS A COMBINATION OF VARIOUS ZERT 6 AND ZERT7 PROGRAMS
C IT USES EITHER THE Z6 OR Z7 COMPUTATION CONTROLLED BY I1
C TIME, A, + XL PUNCHED OUT EVERY 10TH DAY UNDER CONTROL OF I2
C
C D A N G F R D A N G E R
C
C NEVER PUT COMMON /MAN/ IN MAIN PROGRAM
C
C WRITTEN BY CHAN CROCKER MIT LINCOLN LABORATORY
C ROOM C-266 LEXINGTON MASS. TEL. 617-862-5500 EXT. 5307
C COMMENT CARD FPSIL = - 6.* J22
C

```

ISN 0002	C	ALPHABETIZED LIST OF SUBROUTINES	0053
ISN 0003	C	ALVISE IDDAY,VALUEI	0054
	C	CHAN(XLAMI	0055
	C	CURT(IDDAY,VALUEI	0056
	C	GETIMP	0057
	C	MINMAX (X,N,XMIN,XMAXI	0058
	C	NEWFOR (XLIMD,XTEMP,CDNI	0059
	C	SETDUT (PUNCHI	0060
	C	SSL	0061
	C	TACK IXLAMI	0062
	C		0063
	C		0064
	C		0065
	C		0066
	C		0067
	C		0068
	C		0069
	C		0070
	C		0071
	C		0072
	C		0073
	C		0074
	C		0075
	C		0076
	C		0077
	C		0078
	C		0079
	C		0080
	C		0081
	C		0082
	C		0083
	C		0084
	C		0085
	C		0086
	C		0087
	C		0088
	C		0089
	C		0090
	C		0091
	C		0092
	C		0093
	C		0094
	C		0095
	C		0096
	C		0097
	C		0098
	C		0099
	C		0100
	C		0101
	C		0102

ISN 0002	C	IMPLICIT REAL*8 IA-H,O-YI	
ISN 0003	C	DIMENSION RTI15001, TH115001, XL0(1500)	
	C	WARNING.... VARIABLES IN DIFFERENT COMMONS MUST HAVE DIFFERENT NAM	
ISN 0004	C	COMMON/Z8/XLIM, CONST, XLIMX, CUT, XNUM, FUEL,	
	C	1 XX, WS, COSI, ESJN, SINI, TEL, JFK,	
	C	2 LLL, M, K, NUM, I3, I5	
ISN 0005	C	COMMON/PARAMS/FITHDY,CDNSTI, XLIMXD, XOISD, XMASS, CUT3Y, A,	
	C	1 E, XI, OL, XLAMD, TEND, ALJNGR, DATF,	
	C	2 FORCCT, FORLAM, A1, A2, W, I1, I2,	
	C	3 I3T, I4, I5T, I6, I7, NJMIN	
ISN 0006	C	COMMON/SUPSTF/ AB1, AB2, OFFSET, I8	
ISN 0007	C	COMMON/DDVE/XLIMD	
ISN 0008	C	COMMON/EL/CAVE,J1,J2,J3,JANE,KAY,LAD	
ISN 0009	C	EQUIVALENCE IXISP,XDISP)	
ISN 0010	C	DATA G,ZAP,GM,DPRINT,DT,RE /9.800,-1.00,3.996267705,86403.00,	
ISN 0011	C	* 86.400,6378.38800 /	
	C	DATA J20,XJ30,XJ31,XJ32,XJ33/1082.430-6,-2.560-5,-1.40-5,	
	C	I -.1020-6,-.1710-6/	
ISN 0012	C	XX = 0.17453292519943290-D1	
ISN 0013	C	W = .72921151466715650-4	

ISN 0014	ESUN = 0.01672300					0103
ISN 0015	WS = 282.3728700 * XX					0104
ISN 0016	SIN1= OSIN (23.4435800 * XX)					0105
ISN 0017	COS1 = OCOS(23.4435800 * XX)					0106
ISN 0018	XDT = OT					0107
ISN 0019	NDUT = 6.283185300 / (W * DT)					0108
ISN 0020	7 CALL GETINP					0109
ISN 0021	IF (NUMIN .LT. 0) GO TO 10					0110
	C					0111
ISN 0023	J1=1					
ISN 0024	J2=-1					
ISN 0025	J3=1					0115
ISN 0026	KAY = D					
ISN 0027	JANE=1					
ISN 0028	LAD=5					
ISN 0029	CAVE=CONST/32					
	C					
ISN 0030	C MOVE THE ELEMENTS FROM THE NAME0 COMMON TO BLANK COMMON					0119
ISN 0031	NUM = NUMIN					0120
ISN 0032	CONST = CONST					0121
ISN 0033	I3 = I3T					0122
ISN 0034	I5 = I5T					0123
ISN 0035	CUT = CUT8Y					0124
	TWIDY = FITHDY*DPRI					0125
	C					0126
ISN 0036	WRITE (6, 9)					0127
ISN 0037	9 FORMAT(126H DT TOAYS LAM R DR OLAM E. LONG					0129
	1					0130
	2					0131
ISN 0038	NYR = TENO					0132
ISN 0039	IF(12.EQ. 1) WRITE(7,46) NYR					0133
ISN 0041	46 FORMAT (15)					0134
	C					0135
	C					0136
	C					0137
	C					0138
ISN 0042	GO TO (27, 28), I4					0139
ISN 0043	27 XJ22 = -1.8160-6					0140
ISN 0044	GO TO 29					0141
ISN 0045	28 XJ22 = -1.70-6					0142
ISN 0046	XLAM22 = -19.00 * XX					0143
	C					0144
ISN 0047	29 CONTINUE					0145
ISN 0048	XLAM31 = -169.00 * XX					0146
ISN 0049	XLAM32= D.00					0147
ISN 0050	XLAM33 = 24.9 * XX					0148
ISN 0051	EPS1L = -6.00 * XJ22					0149
ISN 0052	XJ2 = XJ20					0150
	C					0151
						0152

```

ISN 0053 C3 = GM * EPSIL * RE * RE
ISN 0054 C1 = (C3/EPISL) * XJ2 * 3.00/2.00
ISN 0055 C2 = C3 * 3.00/2.00
ISN 0056 C4 = C3 / 2.00

C

ISN 0057 M=0
ISN 0058 K=-1
ISN 0059 JFK=0
ISN 0060 N = 0
ISN 0061 XNUM, = NUM
ISN 0062 XNUM = XNUM * 1000.00 * XMASS * XOT / (XISP * G)
ISN 0063 FUEL = 0.00
ISN 0064 XLAM22 = -15.400 * XX
ISN 0065 XLIMX = XLIMXO*XX
ISN 0066 T = OATE * 86400.00
ISN 0067 TPRINT= T + 86400.00
ISN 0068 TOUT = TWTY0Y + T
ISN 0070 V = (90.00-OL)*XX
ISN 0071 P = A * (1.00 - E*E)
ISN 0072 R = P / (1.00 + E * OCOS(V) )
ISN 0073 H = OSQRT(GM*P)
ISN 0074 OR = H * E * OSIN(V)/
ISN 0075 THETA = 90.00 - XI
ISN 0076 THETA = THETA*XX
ISN 0077 TENO = TENO * 86400.00
ISN 0078 XLOEY = XLAMO
ISN 0079 XLAM = XLAMO * XX
ISN 0080 OFFSET = OFFSET * XX
ISN 0081 OTHETA= 0.00

C
C CALCULATE OXLAM
C O2R = H*H*(1.00-R/P)/R**3
C BETTER CALCULATION OF OXLAM O2R IS NOT 0
C GO TO (2, 52),11

C
2 SQ = GM/(R**3 * (OSIN(THETA))**2)
1 -3.00*GM*XJ2*RE*RE*
2 (3.00*(OCOS(THETA))**2-1.00)/(2.00*R**5*(OSIN(THETA))**2)
3 +3.00 * C3 *OCOS(2.00*(XLAM-XLAM22)) / (2.00*R**5)
4 + O2R / (R * (OSIN(THETA))**2)

C
C DXLAM = OSQRT(SQ)-W
C
C CALL SSL
C
1 SITH = OSIN(THETA)
ISN 0088 SITHSO= SITH * SITH

```

```

ISN 0089      COTH = DCOS(THETA)
ISN 0090      COTHSQ= COTH * COTH
ISN 0091      SIXLT2= DSIN(2.00 * (XLAM-XLAM22))
ISN 0092      COXLT2= DCOS(2.00 * (XLAM-XLAM22))
ISN 0093      SITHT2 = OSIN (2.00 * THETA)
ISN 0094      DXLW = OXLAM * W
ISN 0095      DXLSQ= OXLW*DXLW
ISN 0096      R4 = R**4
ISN 0097      R5 = R**5
ISN 0098      TEL = T
ISN 0099      XFORCE = FLAM (XLAM)

CC
ISN 0100      D2R = R * DTHETA**2 + R * SITHSQ * DXLSQ
1             -GM/R**2 + C1 * (3.00 * COTHSQ -1.00) / R4
2             - C2 * SITHSQ * COXLT2
3             + XFORCE * ALONGR
/ R4

CC
ISN 0101      D2XLAM=-C3 * SIXLT2 / R5
1             -DXLW * COTH * 2.00 * DTHETA / SITH
2             -2.00 * DR * DXLW / R
3             + (XFORCE / (SITH * R)) * FORLAM

CC
ISN 0102      D2THET=-2.00 * DR * DTHETA / R
1             +DXLSQ * SITHT2 / 2.00
2             + C1 * SITHT2 / R5
3             + C4 * SITHT2 * COXLT2 / R5 + XFORCE * FORCCT / R
GO TO 23

C
ISN 0103
ISN 0104      S2 SITH = DSIN(THETA)
ISN 0105      SITHSQ= SITH * SITH
ISN 0106      COTH = DCOS(THETA)
ISN 0107      COTHSQ= COTH* COTH
ISN 0108      REDR = RE/R
ISN 0109      REDRSQ= REDRSQ

C
ISN 0110      FR = GM/(R**2) * (-1.00 + REDRSQ *
1             (1.5 * XJ20 * (3.00 * COTHSQ - 1.00)
2             + 9.00 * XJ22 * SITHSQ * DCOS (2.00*(XLAM-XLAM22))
3             + 7.00 * REDR * XJ30 * (5.00 * COTHSQ -3.00)* COTH
4             + 6.00 * REDR * XJ31 * (5.00 * COTHSQ -1.00)* SITH *
5             DCOS(XLAM-XLAM31)
6             +60.00 * REDR * XJ32 * SITHSQ*COTH*DCOS(2.00*(XLAM-XLAM32))
7             +60.00 * REDR * XJ33 * SITH*SITHSQ* DCOS(3.00*
8             (XLAM-XLAM33)
9             ))

C
C BETTER CALCULATION OF DXLAM D2R IS NOT 0
SQ = (D2R -FR) / (R * SITHSQ)
DXLAM = DSQRT(SQ) - W
C
C

```



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0349
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0351
0352

      2      +30.00 * REXJ32      * SICOTH * DSIN(XL32T2)
      3      +45.00 * RXSISQ      * DSIN(XL33T3) )
      4      + XFORCE * FORLAM

C      FTHETA= -GMORSQ * REORSQ * (-3.00 * XJ20 * SICOTH
1      +6.00 * XJ22 * SICOTH * CX22T2
2      -1.500 * REXJ30 * (COTH5-1.00) * SITH
3      +1.500 * REXJ31 * (15.00 * COTH5Q - 11.00) * CJT4 * CXL31
4      +15.00 * REXJ32 * ( 3.00 * COTH5Q - 1.00) * SITH * CX32T2
5      +45.00 * RXSISQ      * COTH      * DCOS(XL33T3) )
6      + XFORCE * FORCCT

C      D2R = R * OTHETA**2 + R * SITHSQ * DXLWSQ + FR
C      D2XLAM= -2.00 * DR * OXLW/R
C      I      -2.00 * COTH * OTHETA * OXLW/SITH + FXLAM/(R*SITH)
C
C      D2THET= -2.00 * DR * OTHETA/R + FTHETA/R + DXLWSQ * SITH2/2.00
C
CHANGE THE FOLLOWING CARDS IF A DIFFERENT OUTPUT MODE IS DESIRED
23 IF (T - TOUT) 21,14, 14
14 IF ( N - NOUT) 15, 15, 16
15 N = N+ 1
      RT(N) = R
      TH(N) = THETA
      XL(N) = XLAM
      GO TO 21
16 CALL MINMAX (RT,N,RP,RA)
      A = (RA +RP) /2.00
      F = RA/A - 1.00
      CALL MINMAX (TH, N, TMIN, TMAX)
      XI = 90.00 - TMIN/XX
      CALL MINMAX (XL,N,XLMIN,XLMAX)
      XLMD = XLMIN/XX
      XLMAO = XLMAX/XX
      DIFFER = XLMAO - XLMD
      WRITE (6 ,25)A,E,XI,XLMD,XLMAO,DIFFER
25 FORMAT(4H A = ,1PE15.6,4H E = ,1PE15.6,5H XI =,1PE15.6,5H LMI=
      ,1PE15.6,5H LMA= ,1PE15.6, 5H DIF= , 1PE15.6)

C      GO TO (12,13),12
12 GO TO (137,138),17
137 Y2 = (XLMD +XLMAO) / 2.00
      GO TO 139
138 Y2 = XLAM/XX
139 CONTINUE
      WRITE (7,114) TDAYS, A, Y2
114 FORMAT (3F20.7)
C

```

ISN 0182	13 TOUT = TOUT + TWYDY	0353
ISN 0183	N = 0	0354
ISN 0184	GO TO 23	0355
ISN 0185	21 IF (T - TPRINT) 4,3,3	0356
	C	0357
ISN 0186	3 TOEG = THETA/XX	0358
ISN 0187	XLOEG=XLAM/XX	0359
ISN 0188	TNAYS = T/86400.00	0360
ISN 0189	DXLDEG= 86400.00 * DXLAM/XX	0361
ISN 0190	OTDEG = 86400.00 * DTHETA/XX	0362
ISN 0191	XLDET = XLOEG	0363
ISN 0192	IF (XLOET-XLOEY) 34, 34, 35	0364
ISN 0193	34 LLL = -1	0365
ISN 0194	GO TO 36	0366
ISN 0195	35 LLL = 1	0367
ISN 0196	36 XLOEY = XLOEG	0368
ISN 0197	XELONG= 360.00 - XLOEG	0369
ISN 0198	XELONG = 360. - XLOEG	0370
	C	0371
ISN 0199	WRITE (6,5)TDAYS,R,DR,TDEG,DTDEG,XLDEG,XELONG	
ISN 0200	5 FURMAT(IX,IPRE15,6)	0373
	C	0374
ISN 0201	K=-1	0375
ISN 0202	TPRINT = TPRINT + OPRIN1	0376
ISN 0203	4 IF (T.LT.TEND) GO TO 6	0377
ISN 0205	IF (I2.EQ.1) WRITE(7,88) ZAP	0378
ISN 0207	88 FORMAT (E20.5)	0379
ISN 0208	GO TO 7	0380
	C	0381
	C	0382
ISN 0209	6 T=FINOV(T,OT1)	0383
ISN 0210	R = OPNV(R,DR)	0384
ISN 0211	THETA = OPNV(THETA,OTHETA)	0385
ISN 0212	XLAM = OPNV(XLAM,DXLAM)	0386
ISN 0213	OR = OPNV(DR,D2R)	0387
ISN 0214	OTHETA = OPNV(OTHETA,O2THET)	0388
ISN 0215	DXLAM = OPNV(DXLAM,O2XLAM)	0389
	C	0390
	C	0391
	C	0392
	C	0393
ISN 0216	THE PROGRAM CAN NEVER GO TO S.N. 10 THIS IS HERE TO PREVENT	
ISN 0217	A DIAGNOSTIC (PROGRAM DOES NOT END WITH A STOP	
ISN 0218	IF(11 -1) 10,1,51	0394
	10 RETURN	0395
	ENO	0396

LEVEL 2 FEB 67

OS/360 FORTRAN H

DATE 68.067/17.15.57

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINFCNT=50,SOURCE,ERCDCIC,NOLIST,NODECK,LOAD,MAP,NODEFIT,TD

```

ISN 0002      SUBROUTINE ALVISE(DDAY,VALUE)
C              SUBROUTINE TO PICK UP A NEW EPHEMERIS CORRECTION VALUE
C              THE ROUTINE ASSUMES TABLES DEFINING A STEP FUNCTION HAVE BEEN CO
C              MPILED AND STORED IN /STEPS/ COMMON
C
C              THE ROUTINE SCANS THE GDAY ARRAY LOOKING FOR THE FIRST DAY LESS
C              THAN THE INPUT DAY. IT THEN USES THE INDEX FOUND BY THE SCAN TO
C              REFERENCE THE CORRECT TABLE
C
ISN 0003      REAL*8 DDAY,VALUE
ISN 0004      DATA DAYOLD,INDOLD,DEGRAD/366.,91.,.174532975E-1/

C              IN AN ATTEMPT TO SPEED UP THE TABLE LOOK UP WE ARE KEEPING TRA
C              OF THE OLD VALUE ON THE LAST CYCLE AND STARTING THE TLJ FROM THE
C              PREVIOUS DAY VALUE
C              COMMON /STEPS/ GDAY(200),CORRECT(200),INDLST
C              DAY = DDAY
C              IF(DAYOLD .GT. DAY) INDOLD = 1
C              IF      = INDOLD + 1
C              DO I=1,INDLST
C              J      = I - 1
C              IF(DAY - GDAY(I)) 2,I,1
C              I CONTINUE
C              J      = INDLST
C              C
C              UNABLE TO FIND A CORRECT VALUE
C              IMPOSSIBLE -----
C
C              ? CONTINUE
C              DAYOLD = GDAY(IJ)
C              INDOLD = J
C              VALUE = CORRECT(J)* DEGRAD
C              RETURN
C              ENO
ISN 0015
ISN 0016
ISN 0017
ISN 0018
ISN 0019
ISN 0020

```

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,EBCDIC,NOLIST,NODECK,LOAD,MAP,NODEFIT,IN

```

ISN 0002      BLOCK DATA
C             DATA SUBROUTINE TO DEFINE THE STEP FUNCTIONS FOR AN EPHEMERIS CO
C
ISN 0003      COMMON /STEPS/ GOAY(200),CORRECT(200),INDLST
C
C             STATE 0
C             DATA GOAY( 1),CORRECT( 1) /0.,.25
C             DATA GOAY( 2),CORRECT( 2) /3.,.50
C             DATA GOAY( 3),CORRECT( 3) /6.,.75
C             DATA GOAY( 4),CORRECT( 4) /9.,1.
C             DATA GOAY( 5),CORRECT( 5) /12.,1.25
C             DATA GOAY( 6),CORRECT( 6) /15.,1.50
C             DATA GOAY( 7),CORRECT( 7) /18.,1.75
C             DATA GOAY( 8),CORRECT( 8) /21.,2.00
C             DATA GOAY( 9),CORRECT( 9) /24.,2.25
C             DATA GOAY(10),CORRECT(10) /27.,2.50
C             DATA GOAY(11),CORRECT(11) /30.,2.75
C             DATA GOAY(12),CORRECT(12) /33.,3.00
C             DATA GOAY(13),CORRECT(13) /36.,3.25
C             DATA GOAY(14),CORRECT(14) /39.,3.50
C             DATA GOAY(15),CORRECT(15) /42.,3.75
C
C             STATE A
C             DATA GOAY(16),CORRECT(16) /45.,3.50
C             DATA GOAY(17),CORRECT(17) /50.,3.25
C             DATA GOAY(18),CORRECT(18) /55.,3.00
C             DATA GOAY(19),CORRECT(19) /60.,2.75
C             DATA GOAY(20),CORRECT(20) /65.,2.50
C             DATA GOAY(21),CORRECT(21) /70.,2.25
C             DATA GOAY(22),CORRECT(22) /75.,2.00
C             DATA GOAY(23),CORRECT(23) /80.,1.75
C             DATA GOAY(24),CORRECT(24) /85.,1.50
C             DATA GOAY(25),CORRECT(25) /90.,1.25
C             DATA GOAY(26),CORRECT(26) /95.,1.00
C             DATA GOAY(27),CORRECT(27) /100.,0.75
C             DATA GOAY(28),CORRECT(28) /105.,0.50
C             DATA GOAY(29),CORRECT(29) /110.,0.25
C             DATA GOAY(30),CORRECT(30) /115.,0.00
C             DATA GOAY(31),CORRECT(31) /120.,-0.25

```

ISN 0035	DATA	GOAY(32),CORECT(32)	/125.,-0.50	/	0480
ISN 0036	DATA	GOAY(33),CORECT(33)	/130.,-0.75	/	0481
ISN 0037	DATA	GOAY(34),CORECT(34)	/135.,-1.00	/	0482
	C				0483
	C				0484
	C				0485
	C				0486
	C				0487
	C				0488
	C				0489
	C				0490
ISN 0038	DATA	GOAY(35),CORECT(35)	/142.,-0.75	/	0491
ISN 0039	DATA	GOAY(36),CORECT(36)	/149.,-0.50	/	0492
ISN 0040	DATA	GOAY(37),CORECT(37)	/156.,-0.25	/	0493
ISN 0041	DATA	GOAY(38),CORECT(38)	/163.,0.00	/	0494
ISN 0042	DATA	GOAY(39),CORECT(39)	/170.,0.25	/	0495
ISN 0043	DATA	GOAY(40),CORECT(40)	/177.,0.50	/	0496
ISN 0044	DATA	GOAY(41),CORECT(41)	/184.,0.75	/	0497
ISN 0045	DATA	GOAY(42),CORECT(42)	/191.,1.00	/	0498
ISN 0046	DATA	GOAY(43),CORECT(43)	/198.,1.25	/	0499
ISN 0047	DATA	GOAY(44),CORECT(44)	/205.,1.50	/	0500
ISN 0048	DATA	GOAY(45),CORECT(45)	/212.,1.75	/	0501
	C				0502
	C				0503
	C				0504
	C				0505
	C				0506
	C				0507
	C				0508
	C				0509
	C				0510
	C				0511
	C				0512
	C				0513
	C				0514
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	C				0696
	C				0697
	C				0698
	C				0699
	C				0700

COMPILER OPTIONS - NAME= MAIN,DPT=0,LINECNT=50,SOURCE,ERRCIC,NOLIST,NDOECK,LOAD,MAP,NDOFIT,TD

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ISN 0002      SUBROUTINE CURT (DDAY,VALUE)
C             SUBROUTINE TO PICK UP AN EPHEMERIS SUN SENSOR VALUE
C             THE ROUTINE ASSUMES TABLES DEFINING A STEP FUNCTION HAVE BEEN
C             COMPILED AND STORED IN /SPETS/ COMMON
C             THE ROUTINE SCANS THE BODY ARRAY LOOKING FOR THE FIRST DAY LESS
C             THAN THE INPUT DAY. IT THEN USES THE INDEX FOUND BY THE SCAN TO
C             REFERENCE THE BORECT TABLE
C             REAL * 8 DDAY, VALUE
C             DATA DAYOLD, INDOLO /366.,24./
C             AN ATTEMPT TO SPEED UP THE LOOK UP WE ARE KEEPING TRACK OF THE
C             OLD VALUE ON THE LAST CYCLE AND STARTING THE TLJ FROM THE
C             PREVIOUS DAY VALUE.
C             COMMON /SPETS/ BDAY(100),BRECT(100),INDEX
C             DAY = DDAY
C             IF(DAYOLD.GT. DAY) INDOLO = 1
C             II = INDOLO + 1
C             DO 1 I=II,INDEX,1
C             J = I-1
C             IF(DAY-HDAY(I)) 2,1,1
C             1 CONTINUE
C             J = INDEX
C             2 CONTINUE
C             DAYOLD = BDAY(J)
C             INDOLO = J
C             VALUE = BRECT(J)
C             RETURN
C             ENO
ISN 0003
ISN 0004
ISN 0005
ISN 0006
ISN 0007
ISN 0009
ISN 0010
ISN 0011
ISN 0012
ISN 0013
ISN 0014
ISN 0015
ISN 0016
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COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,FBCOIC,NOLIST,NJDECK,LDAD,MAP,N7EDIT,I0
```

[illegible]


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0673 7(XVF(37(1), VFI3(1,7)), (XVFI38(1),VFI3(1,8)),
0674 8 (XVFI39(1),VFI3(1,9)), (XVFI310(1),VFI3(1,10)),
0675 9 (XVFI61(1), VFI6(1,1)), (XVFI60(1),VFI6(1,3)),
0676 2,(XVFI81(1),VFI8(1,1)), (XVFI82(1),VFI8(1,2))
0677 1, (XVFI80(1),VFI8(1,4)) , (XVFI62(1),VFI6(1,2))
0678 3,(XVFI83(1),VFI8(1,3))
0679
0680 R IS ENTERED IN KM/DR IN KM/SEC, THETA IN DEGREES, DTHETA IN
0681 DEGREES PER DAY, XLAM IN DEGREES, DXLAM IN DEGREES PER DAY
0682 TFND IS ENTERED IN SOLAR DAYS
0683 XLAM MEASURED EAST OF GREENWICH
0684 NAV NUMBER OF DAYS TO AVERAGE XTEMP OVER FOR NEWFR
0685
0686 WHEN I1 = 1, Z6 CALC IS USED AND WHEN I1 = 2, Z7 CALC IS USED
0687 WHEN I2 = 1, PUNCH OUTPUT IS EXECUTED))WHEN I2 = 2, NO PUNCH
0688 I3 CONTROLS SUN CORRECTION
0689 I3 = 1 NO CORRECTION
0690 I3 = 2 CORRECTED LINEARLY
0691 I3 = 3 CORRECTED LINEARLY W/ RANDO
0692 I3 = 4 CORRECTED PERFECTLY
0693 I3 = 5 CORRECT BY STEP FUNCTION
0694 I3 = 6 CORRECTED BY S-F W/ RANDOM
0695 I3 = 7 CORRECT BY EPHEMERIS SUN SFN
0696 I3 = 8 CORRECT BY EPHEMERIS SS W/R
0697 I3 = 9 1ST SUN SAIL CORRECTION
0698 I3 = 10 SOLAR SAIL3 FIXED SAIL
0699 I3 = 11 SOLAR SAIL2
0700 IF I4 = 1 XJ22=-1.816E-6 AND XLAM22 = -15.0*.01745
0701 IF I4 = 2 XJ22=-1.70E-6 AND XLAM22 = -19.0*.01745
0702 IF I5 = 1, NO FORCE CUTOFF
0703 IF I5 = 2, FORCE CUT BY INPUT NO. CUT
0704 I6 = 1 SOLAR SAIL2
0705 I6 = 2 SAIL IS TACKED
0706 IF I8 = 1, SAIL IS FLIPPED WITH THE ORBIT
0707 IF I8 = 2, SAIL IS FLIPPED AGAINST THE ORBIT
0708 IF I8 = 3, SAIL IS FLIPPED BY LOGIC IN SATELLITE
0709
0710 CHARACTER TO BE PUNCHED TO SIGNAL THE END OF THE DATA OUTPUT
0711 DATA ZAP/-1./
0712
0713 VARIABLE FORMATS USED FOR THE OUTPUT
0714
0715 DATA XVFI83 /'(45H SAIL IS FLIPPED USING LOGIC IN THE SATELLITE)'/
0716 DATA XVFI82 /'(33H SAIL IS FLIPPED AGAINST THE ORBIT)'/
0717 DATA XVFI81 /'(31H SAIL IS FLIPPED WITH THE ORBIT)'/
0718 DATA XVFI80 /'(10H
0719 DATA XVFI72 /44H(39H PLOTTING XLAM AT END OF EVERY 10TH DAY)/
0720 DATA XVFI71 /43H(38H PLOTTING AVERAGE OF XLAM FOR 10TH DAY)/
0721 DATA XVFI62 /'(12H TACKED SAIL)'/
0722 DATA XVFI61 /25H(20H SOLAR SAIL SCHEME 2)/

```


ISN 0086	AVERAG = .TRUE.	0R21
ISN 0087	CUTDWN = .FALSE.	0R22
ISN 0088	EPH = .FALSE.	0R23
ISN 0089	EPHRAN = .FALSE.	0R24
ISN 0090	FLIPAG = .FALSE.	0R25
ISN 0091	FLIPW = .FALSE.	0R26
ISN 0092	LINEAR = .FALSE.	0R27
ISN 0093	LINRAN = .FALSE.	0R28
ISN 0094	NEWFR=.FALSE.	0R29
ISN 0095	PERFCT = .FALSE.	0R30
ISN 0096	PRINT = .TRUE.	0R31
ISN 0097	PUNCH = .TRUE.	0R32
ISN 0098	SAIL1 = .FALSE.	0R33
ISN 0099	N = 1	0R34
ISN 0100	SAIL2 = .FALSE.	0R35
ISN 0101	SAIL3 = .FALSE.	0R36
ISN 0102	SAIL4 = .FALSE.	0R37
ISN 0103	STEP = .FALSE.	0R38
ISN 0104	STPRAN = .FALSE.	0R39
ISN 0105	TACSAL = .FALSE.	0R40
ISN 0106	THIRD = .FALSE.	0R41
		0R42
		0R43
		0R44
		0R45
		0R47
		0R48
		0R49
		0R46
		0R50
		0R51
		0R52
		0R53
		0R54
		0R55
		0R56
		0R57
		0R58
		0R59
		0R60
		0R61
		0R62
		0R63
		0R64
		0R65
		0R66
		0R67
		0R68

ISN 0107	READ (5,SETUP)	
ISN 0108	C1 = CONST	
ISN 0109	C2 = CONST / 4.00	
ISN 0110	C3 = CONST	
ISN 0111	CONN=CONST/32	
ISN 0112	DMAX=DARS(C1)/25.00	
ISN 0113	XLAM = XLAMD	
ISN 0114	CHECK NIM FOR THE END OF DATA BEFORE WE GO THROUGH THE OUTPUT PHASE OF THIS SECTION IF(NUM .LT. 0) GO TO 10	
	MODIFY THE CONTROL VARIABLES AS A FUNCTION OF THE INPUT	
ISN 0116	IF(THIRD) I1=2	
	DEFINE THE PUNCH VARIABLE FOR THE OUTPUT SUBROUTINE IF(THIRD) I2=1	
ISN 0118	IF(PUNCH) I2 = I	
ISN 0120	IF(I2 .EQ. 1) PUNCH = .TRUE.	
ISN 0122	CALL SETOUT(PUNCH)	
ISN 0123	IF(AVERAG) I7=1	

85

ISN 0185	IF (SAIL4) I6 = 0	0919
ISN 0187	CALL RITCUT (VFI6 (1,16))	0920
ISN 0188	C	0921
	IF (15-1) 1,1,3	0922
ISN 0189	C	0923
ISN 0190	1 WRITE (6,39)	0924
ISN 0192	IF (PUNCH) WRITE (7,39)	0925
	GO TO 4	0926
C		0927
ISN 0193	3 WRITE (6,40) CUT	0928
ISN 0194	IF (PUNCH) WRITE (7,40) CUT	0929
ISN 0196	4 CONTINUE	0930
ISN 0197	CALL RITCUT (VFI7 (1,17))	0931
ISN 0198	CALL RITCUT (VFI9 (1,18))	0932
ISN 0199	WRITE (6,50) IRUN	0933
ISN 0200	IF (PUNCH) WRITE (7,50) IRUN	0934
ISN 0202	IF (SAIL4) I6 = 1	0935
ISN 0204	IF (16.EQ.3) I6 = 0	
ISN 0206	IF (18.EQ.4) I8 = 0	
ISN 0208	RETURN	0936
ISN 0209	C	0937
ISN 0211	10 IF (PUNCH) WRITE (7,11) ZAP,ZAP	0938
	11 FORMAT (2F20.7)	0939
ISN 0212	C	0940
ISN 0213	RETURN	0941
	END	0942

LEVEL 2 FEB 67

OS/360 FORTRAN H

DATE 68.067/17.16.23

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,ERCDIC,NOLIST,NJOECK,LOAD,MAP,NODEFIT,IP

ISN 0002	SUBROUTINE MINMAX (X, N, XMIN, XMAX)	0943
ISN 0003	DIMENSION X(1500)	0944
ISN 0004	DOUBLE PRECISION X, XMIN, XMAX	0945
ISN 0005	XMIN = X(1)	0946
ISN 0006	XMAX = X(1)	0947
ISN 0007	DO 10 I=2,N	0948
ISN 0008	IF (X(I)-XMIN) 1,2,2	0949
ISN 0009	1 XMIN = X(I)	0950
ISN 0010	2 IF (XMAX-X(I)) 3,10,10	0951
ISN 0011	3 XMAX = X(I)	0952
ISN 0012	10 CONTINUE	0953
ISN 0013	RETURN	0954
ISN 0014	END	0955

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,FBCDIC,NOLIST,NJOECK,LOAD,MAP,NOFDIT,LD

```

ISN 0002      SUBROUTINE NEWFOR(XLIMO,XTEMP,CON)
C IN THIS PROGRAM WE WILL CALCULATE A FORCE,CON. THIS FORCE WILL DEPEND
C ON WHICH REGION THE VARIABLE XTEMP IS IN. WHEN WE GET THE SEQUENCE
C OF REGIONS +A-A+A OR -A-A-A WE WILL CALCULATE THE FORCE DIFFERENTLY.
C
ISN 0003      IMPLICIT REAL*8 (A-H,O-Z)
ISN 0004      DIMENSION CCOUNT(1800) , XT(35)
ISN 0005      COMMON/MAN/C1,C2,C3,OMAX,CONN,NAV,N
ISN 0006      COMMON/EL/CAVE,J1,J2,J3,JANE,K,LAO
ISN 0007      DATA NI /0/

ISN 0008      C THIS SUBROUTINE WORKS WITH QUANTITIES IN DEGREES(UNITS)
ISN 0009      XTEMP = XTEMP / .1745329251994329D-01
ISN 0010      NI=NI+1
ISN 0011      XT(NI)=XTEMP
ISN 0012      IF(NI.EQ.1) GO TO 8
ISN 0013      CUN=CONN
ISN 0014      IF(NI.LT.NAV)RETURN
ISN 0015      SUM=0.0
ISN 0016      DO 3 J=1,NAV
ISN 0017      SUM=XT(J) + SUM
ISN 0018      3 SUM=XT(J) + SUM
ISN 0019      XTEMP=SUM/NAV
ISN 0020      NI=0
ISN 0021      8 N=2

ISN 0022      C THIS SECTION WILL DETERMINE WHERE XTEMP LIES
ISN 0023      IF(XTEMP .GE.XLIMO) GO TO 10
ISN 0024      GO TO 20

ISN 0025      C AT THIS POINT XTEMP IS GREATER THAN OR EQUAL TO XLIM
ISN 0026      10 R=-1.0D0
ISN 0027      I=-1
ISN 0028      IF(XTEMP.GE.(XLIMO+4.00))GO TO 30
ISN 0029      IF(XTEMP.GE.(XLIMO+2.00))GO TO 50
ISN 0030      GO TO 70

ISN 0031      C AT THIS POINT XTEMP IS LESS THAN XLIMO
ISN 0032      20 R=1.0D0
ISN 0033      I=1
ISN 0034      IF(XTEMP.LE.(XLIMO-4.00))GO TO 30
ISN 0035      IF(XTEMP.LE.(XLIMO-2.00))GO TO 50
ISN 0036      GO TO 70

ISN 0037      C XTEMP IS IN REGION C. SET J1=3 OR -3, FORCE=C3 OR -C3 AND J2,J3=0.
ISN 0038      C TEST TO SEE IF WE HAVE BEEN IN C AWHILE
ISN 0039      30 INEW = I*3
ISN 0040      IF(INEW .EQ. J1)GO TO 600
ISN 0041      J1 = INEW
ISN 0042      CONN = R*C3
ISN 0043      K = I
ISN 0044

```

ISN 0045	COUNT(K) = CONN	0993
ISN 0046	J2 = 0	0994
ISN 0047	J3 = 0	0995
ISN 0048	JANE = 0	0996
ISN 0049	CAVE = 0.00	0997
ISN 0050	LAO = 0	0998
ISN 0051	GO TO 600	0999
		1000
		1001
		1002
		1003
		1004
		1005
		1006
		1007
		1008
		1009
		1010
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ISN 0052	C	
ISN 0053	C	WE ARE IN REGION B. HAVE WE CHANGED REGIONS--IF SO, UPDATE AND SET
ISN 0055	C	FORCE = C2 OR -C2.
ISN 0057	50	INew = I*2
ISN 0058		IF(INew.EQ. J1) GO TO 600
		IF(J1.EQ.0) K=0
		CONN = R*C2
		GO TO 205
	C	
	C	WE ARE IN REGION A, AND MUST DECIDE WHAT VALUE TO GIVE THE FORCE.
	C	CHECK TO SEE IF WE ARE IN THE SAME REGION. HAVE WE CHANGED REGIONS?
	70	INew = I*1
		IF(J1.EQ.0) K=0
		IF(INew.EQ. J1) GO TO 600
	C	DIO WE JUST GO FROM B TO A.
		IF(IABS(J1).EQ.2) GO TO 205
	C	TEST FOR INITIAL SEQUENCE
		IF(JANE.EQ.0) GO TO 180
	C	
	C	HAVE PASSED THE INITIAL SEQUENCE AND CHANGED REGIONS. TEST =J1 MAX DA
	C	AND ZERO BUMP SECTION.
	170	IF(IABS(INew*J1*J2).NE.1.AND. LAO.EQ.0) GO TO 190
		IF(IABS(INew*J1*J2).NE.1.AND. LAO.EQ.5) GO TO 181
		IF(OABS(CONN).NE.0.00.AND. OABS(CONN).LT.0MAX) GO TO 200
		IF(LAO.EQ.5) GO TO 5
	C	
	C	IN A AND SEQUENCING. JUST CHANGED REGIONS. MAX DAMP AND ZERO BUMP TAK
	C	TAKEN CARE OF. ALSO COMING FROM B. (INew*J1*J2=1)
	110	IF(OABS(COUNT(K)).NE. OABS(COUNT(K-1))) GO TO 135
	C	
	C	IF WE GET HERE IT MEANS WE ARE READY TO DIVIDE BY 2
	130	CONN = -CONN/2.00
		CAVE = CONN
		IF(OABS(CONN).LT.0MAX) LAO=5
		GO TO 205
	C	
	C	ABSOLUTE VALUE REMAINS THE SAME IF J3.NF.2.
	135	IF(IABS(J3).EQ.2) GO TO 130
		CONN = -CONN
		CAVE = CONN
		GO TO 205
	C	
	C	TEST HERE FOR INITIAL SEQUENCE.

ISN 0088	130 IF(IABS(INEW#J1*J2).EQ.1)GO TO 207	1043
ISN 0090	C WE ARE IN A BUT HAVE NOT REACHED INITIAL SEQUENCE.	1044
ISN 0091	CONN = R*C1	1045
	GO TO 205	1046
	C	1047
ISN 0092	C IN MAX DAMP ZERO BUMP SECTION AND INEW#J1*J2 NE 1	1048
ISN 0094	181 IF(CONN.EQ.0.00)GO TO 5	1049
	GO TO 200	1050
	C	1051
ISN 0095	C INITIAL SEQUENCE JUST REACHED. MEANS CHANGE OF REGION .	1052
ISN 0096	207 JANE = 1	1053
ISN 0097	CONN = R*C1/2.00	1054
ISN 0098	CAVE = CONN	1055
	GO TO 205	1056
	C	1057
ISN 0099	C IN A AFTER THE INITIAL SEQUENCE. INEW#J1*J2 .NE. 1 .	1058
ISN 0100	190 CONN=-CAVE	1060
	GO TO 205	1061
	C	1062
ISN 0101	C IN ZERO BUMP AND MAXIMUM DAMP SECTION	1063
ISN 0102	5 CONN = R*C1/32.00	1064
	GO TO 205	1065
	C	1066
ISN 0103	C FROM B TO A	1067
	200 CONN = 0.000	1068
ISN 0104	C REGIONS HAVE JUST CHANGED. UPDATE SEQUENCE	1069
ISN 0105	205 K = K+1	1070
ISN 0106	J3 = J2	1071
ISN 0107	J2 = J1	1072
ISN 0108	J1 = INEW	1073
	COUNT(K) = CONN	1074
ISN 0109	C	1075
ISN 0110	600 CON = CONN	1076
ISN 0111	RETURN	1077
	END	

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LEVEL 2 FEB 67                                OS/360  FORTRAN H                                DATE  68.067/17.16.32

      COMPILER OPTIONS - NAME=  MAIN,OPT=00,LINECNT=50, SOURCE,EBCDIC,NOLIST,NODECK,LOAD,MAP,NODEIT,ID

      ISN 0002      SUBROUTINE  SETOUT(PUNCH)                                1078
      C              SUBROUTINE  TO OUTPUT ON LUN 6 AND THE PUNCH TAPE  LUN 7    1079
      C              THE VARIABLE FORMAT INFORMATION                          1080
      ISN 0003      LOGICAL PUNCH,CARDS                                     1081
      C                                                      1082
      C              ENTRY POINT TO DEFINE THE LOGICAL VARIABLE CARDS FOR THE SECCNO 1083
      C              ENTRY POINT  RITOUT                                       1084
      ISN 0004      CARDS = PUNCH                                           1085
      ISN 0005      RETURN                                                  1086
      C                                                      1087
      C * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * *  1088
      C * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * * *  1089
      C              ENTRY POINT TO DO THE ACTUAL WRITE AND PUNCH             1090
      C                                                      1091
      ISN 0006      ENTRY RITOUT (V)                                         1091
      ISN 0007      DIMENSION V(18)                                          1092
      ISN 0008      WRITE(6,V)                                               1093
      ISN 0009      IF(CARDS)  WRITE(7,V)                                     1094
      ISN 0010      RETURN                                                  1095
      ISN 0011      END                                                    1096
      ISN 0012

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LEVEL 2 FEB 67	DS/360	FORTRAN H	DATE 58.067/17.16.35
- COMPILER OPTIONS - NAME= MAIN,DPT=00,LINECNT=50, SOURCE,EBCDIC,NOLIST,NJOECK,LOAD,MAP,NODEIT,LD			
ISN 0002	SUBROUTINE SSL		1098
ISN 0003	DOUBLE PRECISION X		1099
ISN 0004	DOUBLE PRECISION CIND		1100
ISN 0005	X=CIND(0,X,0.000)		1101
ISN 0006	RETURN		1102
ISN 0007	END		1103

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,FBC01C,NOLIST,NJOECK,LOAD,MAP,NOFDIT,LD

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1SN 0002      C      SUBROUTINE TACK(XLAM)
C
1SN 0003      C      WRITTEN BY ELLEN SWENSON X5309
1SN 0004      C      IMPLICIT REAL * 8 (A-H,O-Z)
1SN 0005      C      DOUBLE PRECISION LORSTR
1SN 0006      C      LOGICAL CHECK
C              CHECK = .FALSE.
C
1SN 0007      C      COMMON/28/XLIM , CONST , XLIMX , CUT , XNUM , FUEL ,
1          XX , WS , COSI , FSUN , SINI , TEL , JFK ,
2          LLL , M , K , NUM , I3 , I5
C      COMMON/PARAMS/FITHDY,CONSTI, XLIMXD, XDISP, XMASS , CUTBY , A ,
1          E , XI , OL , XLAMD , TENO , ALONGR , DATE ,
2          FORCCT,FORLAM, A1 , A2 , W , I1 , I2 ,
3          I3T , I4 , I5T , I6 , I7 , NUMIN
C
1SN 0009      C      COMMON/SUPSTF/ AB1, AR2, OFFSET, I8
C      OBJECT TO OBTAIN MAXIMUM FORCEE ALONG ORBITAL TANGENT
C
C      UNDERLYING THEORY
C      LARRY BLACK'S IOEA-- OPTIMUM POSITION FOR SAIL IS ALONG
C      BISECTOR OF THE ANGLE BETWEEN THE PROJECTION OF THE SUN IN T
C      ORBITAL PLANE AND THE ORBITAL TANGENT
C      THERE IS PROVISION FOR CHANGING THE POSITION OF SAIL
C      ROTATE SAIL *OFSET*DEGREES FROM TACKED POSITION
C      THERE IS AN ENTRY CHAN AND I8 VARIABLES
C      CHAN PLACES SAIL ON THE RADIUS
C      IR FLIPS THE SAIL THREE WAYS
C      FLIPS SAIL WITH ORBIT
C      FLIPS SAIL AGAINST ORBIT
C      FLIPS SAIL EVERYWHERE
C
C      BOTH SIDES OF THE SAIL ARE REFLECTIVE
C      AB1 = .100
C      AR2 = .100
C
1SN 0010      C      ALONGR IS FORCE ALONG R
1SN 0011      C      FORCCT IS FORCE ALONG CTHEIA
C      EORLAM IS EORCE ALONG LAMDA
C
C      OARCUS(DOTPRO)IS ANGLE BETWEEN NORMAL TO THE SAIL AND THE SUN
C      CTHEIA IS ANGLE BETWEEN ARIES AND RADIUS
C
1SN 0012      C      300 CONTINUE
C
C      GEOCENTRIC CORRDINATES OF NORMAL TO RADIUS
1SN 0013      C      I CTHEIA = XLAM + W * TEL
1SN 0014      C      XNORMA = DSIN(CTHEIA)

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ISN 0015      YNORMA = -OCOS(CTHETA)
ISN 0016      ZNORMA = 0.000

C
C   GEOCENTRIC COORDINATES OF SUN
C   V = TRUE ANOMALY (MOULYON P171)      XM = MEAN ANOMALY
      DAY = TEL/86400.00
      2 XM = (.98560000 * DAY -3.392900) * XX
      1 V = XM + 2.00 * ESUN *OSIN(XM) + 5.00 * ESUN *ESUN * OSIN(2.00
        * XM) / 4.00

C
      U = V + WS
      SINU= OSIN(U)

C
      XSUN= OCOS(U)
      YSUN= SINU * COSI
      ZSUN= SINU * SINI

C   GEOCENTRIC COORDINATES OF THE ORBITAL TANGENT
      XFORCE = -XNORMA
      YFORCE = -YNORMA
      ZFORCE = ZNORMA

C   GEOCENTRIC COORDINATES OF PROJECTION OF THE SUN IN THE ORBITAL PLNE
      XMAINE =OSQRT( XSUN**2 + YSUN**2 )
      XPROJ = XSUN/XMAINE
      YPROJ = YSUN/XMAINE

C   IF (CHECK) GO TO 310

C   COORDINATES OF SAIL ITSELF
      XTEMP = XFORCE + XPROJ
      YTEMP = YFORCE + YPROJ

C   COORDINATES OF NORMALIZED SAIL VECTOR
      YMAINE= DSQRT( XTEMP**2 + YTEMP**2)
      XTEMP = XTEMP / YMAINE
      YTEMP = YTEMP / YMAINE

C   COORDINATES OF SAIL ITSELF
      SINDOFF= OSIN(OFFSET)
      COSOFF= OCOS(OFFSET)
      XRESAL= XTEMP* COSOFF + YTEMP*SINDOFF
      YRESAL= YTEMP * COSOFF - XTEMP * SINDOFF
      XSAIL = YRESAL
      YSAIL = -XRESAL
      GO TO 3

C
      310 CONTINUE
      XSAIL = YNORMA
      YSAIL = YNORMA

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ISN 0048      XRESAL = -YSAIL
ISN 0049      YRESAL = XSAIL

ISN 0050      3  00TPRO = XSAIL *XSUN + YSAIL * YSUN

C
C
C  ANGLE BETWEEN SAIL AND RADIUS = ANCHOR
  LOBSTRE = XFORCE*(-XSAIL) + YFORCE*(-YSAIL)
  ANCHOR = DARCOS(LOBSTR)
  BOAT = DSIN(ANCHOR)

C
  ROCK = OARS(LOBSTR)
  TRAP = DABS(00TPRO)
  IF(00TPRO .LE. 0.00) AB1 = AB2
  PUT = (2.000 - AB1) * 00TPRO**2

C
C  ALUSAL IS FORCE ALONG SAIL DUE TO NON-PERFECT SAIL
  ALUSAL = - AB1 * TRAP * (XSUN*XRESAL + YSUN*YRESAL)

C
C  TO DETERMINE DIRECTION OF FORCE COMPONENTS OF ALONGR AND FORLAW
  IF (18 .EQ. 1 .AND. 00TPRO .LE. 0.00) GO TO 40
  IF (18 .EQ. 2 .AND. 00TPRO .GE. 0.00) GO TO 40
  IF (18 .EQ. 3) GO TO 40

C
C 000 FISH HERRING SALT
C ARE INITIALIZED TO +1 AND MADE NEGATIVE WHEN NECESSARY
C
C 000 IS NEGATIVE WHEN
C DIRECTION OF SOLAR FORCE IS MORE THAN 90 DEGS FROM RADIUS
C
C FISH IS NEGATIVE WHEN
C SAIL IS MORE THAN 90 DEGREES FROM RADIUS
C
C HERRING IS NEGATIVE WHEN
C DIRECTION OF SOLAR FORCE IS MORE THAN 90DEG FROM ORBITAL TAN
C
C SALT IS NEGATIVE WHEN
C SAIL IS MORE THAN 90 DEGREES FROM ORBITAL TANGENT

C
C 000 = 1.000
C FISH = 1.000
C HERRING = 1.000
C SALT = 1.000

C
C COORDINATES OF RADIUS
C X COORDINATE -YNORMA
C Y COORDINATE XNORMA
C
C COORDINATES OF SOLAR RADIATION FORCE
C TESTING ON Z-COMPONENT OF CROSS PRODUCT OF SAIL CROSS SUN
C

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ISN 0070	IF((XRESAL * YPROJ - XPROJ * YRESAL) .LE.0.00) GO TO 7	1253
ISN 0072	XSORAF= XSAIL	1254
ISN 0073	YSORAF= YSAIL	1255
ISN 0074	GO TO 6	1256
C		1257
ISN 0075	7 XSORAF= -XSAIL	1258
ISN 0076	YSORAF= -YSAIL	1259
C		1260
ISN 0077	6 CONTINUE	1261
ISN 0078	IF((XSORAF*XFORCE + YSORAF*YFORCE).LE.0.000) HERING=-1.000	1262
ISN 0080	IF((XPESAL * XFORCE +YRESAL * YFORCE) .LE.0.00) SALT=-1.00	1263
ISN 0082	IF((XSORAF * (-YNORMA) + YSORAF * XNORMA).LE.0.000) C00=-1.000	1264
ISN 0084	IF((-YNORMA)*XRESAL + XNORMA * YRESAL .LE. 0.00) FIST=-1.00	1265
C		1266
C		1267
ISN 0086	FORLAM= HERING*POT*POCK + SALT*ALOSAL*BOAT	1268
ISN 0087	ALONGR= C00*POT*BOAT + FISH*ALOSAL*ROCK	1269
ISN 0088	FORCCT= -AB1 * TRAP* ZSUN	1270
C		1271
ISN 0089	RETURN	1272
C		1273
ISN 0090	40 CONTINUE	1274
ISN 0091	DOTPR2= DABS(ZSUN)	1275
ISN 0092	DOTPR1= XSUN * (-YSAIL) + YSUN * XSAIL	1276
ISN 0093	FORLAM = AB1 * DOTPR2 * DOTPRO	1277
ISN 0094	ALONGR = -AB1 * DOTPR2 * DOTPR1	1278
C		1279
C	TO OBTAIN FORCCT TEST ON SIGN OF ZSUN	1280
ISN 0095	IF(ZSUN .GE.0.00) FORCCT = -(2.00 - AB1)*DOTPR2**2	1281
ISN 0097	IF(ZSUN .LE.0.00) FORCCT = (2.00 - AB1)*DOTPR2**2	1282
ISN 0099	RETURN	1283
C		1284
ISN 0100	ENTRY CHAN(XLAM)	1285
ISN 0101	ROTAAT= XLAM + W*TEL	1286
ISN 0102	CHECK = .TRUE.	1287
ISN 0103	GO TO 300	1288
ISN 0104	END	1289

1339		DY1(I)=DY		1339
1340		GO TO 3000		1340
1341		YO(I)=YO(I)+H*(A3*DY+A4*DY1(I)-A5*DY0(I))		1341
1342	1600	CIND=YO(I)		1342
1343		GO TO 3000		1343
1344	1700	CIND=YO(I)+H*(A6*DY-A7*DY1(I)+A8*DY0(I))		1344
1345		GO TO 2000		1345
1346	1800	CIND=YO(I)+H*(A9*DY+A10*DY2(I)-A11*DY1(I)+A12*DY0(I))		1346
1347		GO TO 2000		1347
1348	1900	CIND=Y+H*(A1*DY-DY2(I)*A2+A3*DY1(I)-A4*DY0(I))		1348
1349	2000	DY0(I)=DY1(I)		1349
1350	2100	DY1(I)=DY2(I)		1350
1351	2200	DY2(I)=DY		1351
1352	3000	IF (I-50) 4000,5000,5000		1352
1353	4000	I=I+1		1353
1354		RETURN		1354
1355	5000	STOP		1355
1356		END		1356

[illegible]

LEVEL 2 FEB 67	OS/360 FORTRAN H	DATE 68.067/17.16.56
<pre> COMPILER OPTIONS - NAME= MAIN,OPT=00,LINENCT=50,SOURCE,EPGOTIC,NOLIST,NODECK,LOAD,MAP,NODEDIT,TD </pre>		
ISN 0002	DOUBLE PRECISION FUNCTION DPNV(Y,DY)	1438
ISN 0003	DOUBLE PRECISION Y,DY	1439
ISN 0004	DOUBLE PRECISION CIND	1440
ISN 0005	DPNV=CIND(1,Y,DY)	1441
ISN 0006	RETURN	1442
ISN 0007	END	1443

LEVEL 2	FER 67	OS/360	FORTRAN H	DATE	58.067/17.16.59
COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCE,ERCNIC,NOLIST,LOAD,WAP,NOCFIT,ID					
15N 0002	DOUBLE PRECISION FUNCTION FINDV(X,H)				1444
15N 0003	DOUBLE PRECISION X,H				1445
15N 0004	DOUBLE PRECISION CIND				1446
15N 0005	FINDV=CIND(0,X,H)				1447
15N 0006	RETURN				1448
15N 0007	END				1449

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50,SOURCEF,EBCDIC,VOLIST,NBDECK,LOAD,MAP,NDEDIT,10

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1450      C      DOUBLE PRECISION FUNCTION FLAM (XLAM)
1451      C
1452      C      THIS FUNCTION COMPUTES A FORCE WHEN THE INPUT PARAMETER IS OUTSID
1453      C      OF A SPECIFIED BAND.
1454      C
1455      C      THE FORCE IS LEFT ON FOR M+1 TIME INTERVALS --- THESE ARE
1456      C      DETERMINED BY THE MAIN PROGRAM AND SET IN COMMON
1457      C
1458      C      THE IMPULSES ARE ALLOWED ON ONE DAY INTERVALS, BUT ONCE STARTED
1459      C      NUM IMPULSES WILL BE GIVEN.
1460      C
1461      C      K IS A GATE WHICH ALLOWS PASSAGE INTO THE ROUTINE. IT IS
1462      C      SET POSITIVE
1463      C      USUALLY POSITIVE, BUT CHANGED TO NEGATIVE ONCE PER DAY.
1464      C      IF IMPULSES ARE BEING GIVEN IT IS LEFT NEGATIVE UNTIL
1465      C      ALL HAVE BEEN COMPLETED.
1466      C
1467      C      XLAM CURRENT POSITION
1468      C
1469      C      XLIM BOTTOM OF THE BAND
1470      C
1471      C      XLIMX TOP OF THE BAND
1472      C      THUS, XLAM-XLIM-X.GT. 0 ---XLAM IS ABOVE THE BAND
1473      C      AND XLAM-XLIM-X.LT. 0 ---XLAM IS BELOW THE BAND
1474      C
1475      C      LL IS CONTROLLED BY THE FORCE AND BY THE POSITION RELATIVE TO
1476      C      THE BAND. IF LL IS NEGATIVE WE ARE INSIDE THE BAND OR TH
1477      C      FORCE HAS BEEN COMPLETED.
1478      C
1479      C      M COUNTS THE NUMBER OF IMPULSES BEING ISSUED
1480      C
1481      C      IMPLICIT REAL * 8 (A-H,O-Z)
1482      C
1483      C      LOGICAL NEWFR
1484      C
1485      C      COMMON/FSL/NEWFR
1486      C      COMMON/DOME/XLIMD
1487      C
1488      C      COMMON/Z8/XLIM , CONST , XLIMX , CUT , XNUM , FUEL ,
1489      C      1 XX , WS , COSI , ESUN , SINI , TEL , JFK ,
1490      C      2 LLL , M , K , NUM , I3 , I5
1491      C
1492      C      COMMON/PARAMS/FITHD,CONSTI, XLIMXD, XDISP, XMASS , CUTBY , A
1493      C      F , XI , J1 , J2 , XLAMD , TEND , ALONGR , DATE ,
1494      C      1 FORCCT,FORLAM, A1 , A2 , W , I1 , I2 ,
1495      C      2 I3T , I4 , I5T , I6 , I7 , NUMIN ,
1496      C      3 COMMON/SUPSTF/ ABI , AB2 , OFFSET , I9
1497      C
1498      C

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C      MODIFY THE INPUT AS A FUNCTION OF THE CORRECTION
C      AND DETERMINE WHERE IT LIES RELATIVE TO THE BAND
C
1549  IF(13.NE.4) X = CORR(TEL)
1550  XTEMP = XLAM-X
1551
C      IF WE ARE USING THE OLD FORCE CALCULATIONS JUST CONTINUE. IF NOT,
C      GO TO SUBROUTINE NEWFOR
1552  IF(NEWFOR) GO TO 300
1553  IF(XTEMP.GT.XLIMX) GO TO 11
1554  IF(XTEMP.LT.XLIM) GO TO 1
1555
C      K=1
1556  GO TO 9
1557
C      WE ARE OUTSIDE AND ABOVE, THRUSTING WITH THE ORBIT
C
1558  11  FORCE=-CON
1559  IF (18.NE.0) 18 = 1
1560  OFFSET = -90.000 * XX
1561  A81 = A2
1562  A82 = A1
1563  WE ARE OUTSIDE AND BELOW, THRUSTING AGAINST THE ORBIT
1564  GO TO 4
1565
C      IF WE REACH HERE IT MEANS WE ARE USING THE NEW FORCE CALCULATIONS
C
1566  300  CALL NEWFOR(XLIMD,XTEMP,CON)
1567  1  FORCE = CON
1568  IF(18.NE.0) 18 = 2
1569  OFFSET = 0.00
1570  A81 = A1
1571  A82 = A2
1572  4  M = M + 1
1573
C      IF(16.EQ.2) CALL TACK(XLAM)
1574  43  IF (16.FO.1) CALL CHAN (XLAM)
1575  IF(M.GT. NUM) GO TO 7
1576
C      WE ARE STILL GIVING THE IMPULSES
C
1577  FLAM = FORCE
1578  IF (16.EQ.1) FLAM = OABS (FORCE)
1579  JFK = JFK + I
1580  RETURN
1581
C      PUBLISH THE REPORT OF THE FUEL CONSUMED
C      WE ARE OUTSIDE AND BELOW
C
1582  7  FUEL=FUEL+OABS(CON*NUM)
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IF (I16.NE. 0) FUEL = 0.00
CORD = -X/XX

WRITE(6,10) JFK,CORD,FORCE,FUEL
10  FORMAT(IX,10,'CORD=',1PE15.6,'CON=',1PE15.6,'FUEL USED=',1PE15.6,
1  'KG')

      SET THE GATE TO CLOSED
      K = 1
      M = 0
      GO TO 9
END

```


LEVEL 2 FEB 67

05/360 FORTRAN H

DATE 68.067/17.17.09

COMPILER OPTIONS - NAME= MAIN,OPT=00,LINECNT=50, SOURCE,FRCDIC,NOLIST,NJDECK,LDAD,MAP,NDEFIT,TD

```
ISN 0002      DOUBLE PRECISION FUNCTION ESMOD (X1,X2)
ISN 0003      IMPLICIT REAL*8 (A-H,O-Z)
ISN 0004      IX=IDINT(X1/X2)
ISN 0005      ESMOD=X1-IX*X2
ISN 0006      RETURN
ISN 0007      END
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REFERENCES

1. M. J. Neufeld and B. M. Anzel, "Synchronous Satellite Stationkeeping," AIAA Paper No. 66-304.
2. J. H. Molitor and M. H. Kaplan, "Optimization of Ion Engine Control Systems for Synchronous Satellites," American Institute of Aeronautics and Astronautics, Summer Meeting, Paper No. 63-273.
3. R. R. Allan and G. E. Cook, "The Long-Period Motion of the Plane of a Distant Circular Orbit," Proc. Royal Society 280, p. 97 (7 July 1964).
4. W. L. Black, M. C. Crocker, and E. H. Swenson, "Stationkeeping a 24-Hour Satellite Using Solar Radiation Pressure," to be published in Journal of Spacecraft and Rockets.
5. L. Blitzer, E. M. Boughton, G. Kang and R. M. Page, "Effect of Ellipticity of the Equator on a 24-Hour Nearly Circular Satellite Orbits," J. of Geophysical Research 67, 329 (1962).
6. C. A. Wagner, "The Drift of a 24-Hour Equatorial Satellite Due To An Earth Gravity Field Through 4th Order," NASA TN D-3317 (1966).
7. C. A. Wagner, "Determination of the Ellipticity of the Earth's Equator from Observations on the Drift of the Syncom II Satellite," NASA TN D-2759 (May 1965).
8. Private communication.
9. Private communication.
10. B. J. Moriarty, "Position Error in Station-Keeping Satellite," Technical Note 1966-21, Lincoln Laboratory, M. I. T. (1 April 1966), DDC 633034.
11. Private communication.

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A program has been written to simulate the east-west stationkeeping of a synchronous satellite. Different ways of implementing the thrust sequence of rocket motors and solar sails are discussed.

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